

Astrometry of Long-Period Variable Stars: Specific Problems and Possible Actions

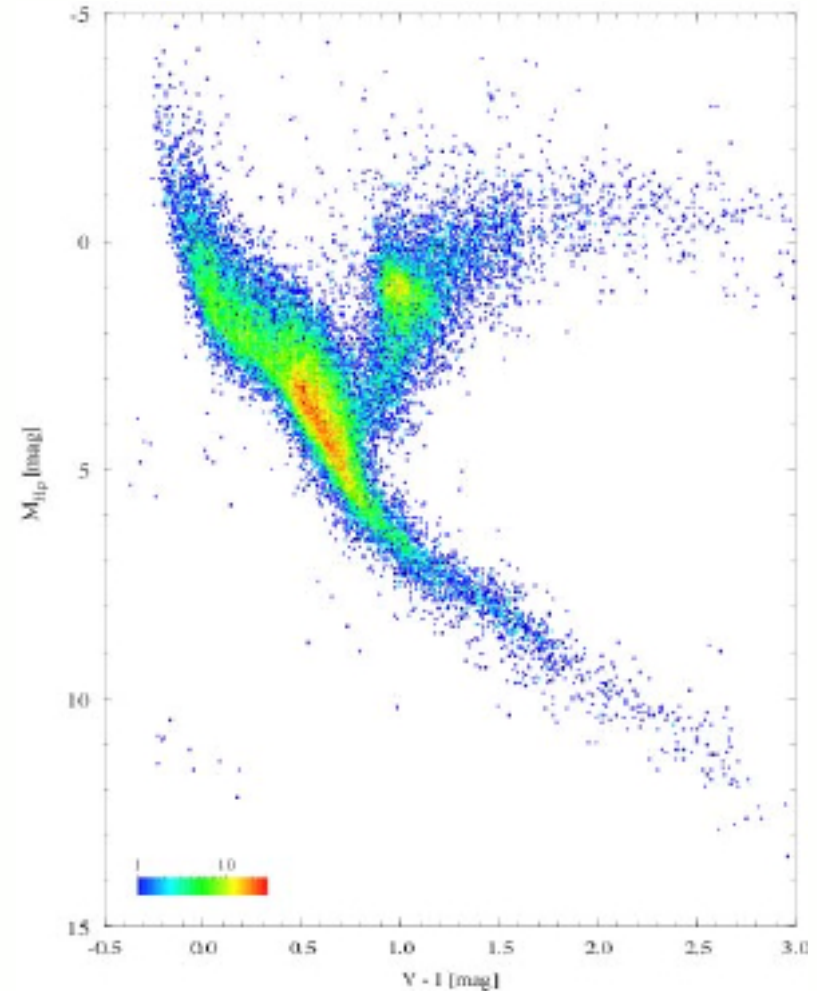
A. Jorissen

Astrometry of Long-Period Variable Stars: Specific Problems

- **Very red colours:**
 - $V-I > 2$ for C stars,**
 - $V-I > 3$ for M stars**

The chromaticity correction is very important for those stars, **albeit difficult to calibrate!**

(because difficult to find stars to calibrate the chromaticity correction in this $V-I$ range, which are not suffering from the problems specific to LPVs)



Astrometry of Long-Period Variable Stars: Specific Problems

- **Variability:**

- Δ V-I as large as 3 mag

- chromaticity correction
is varying over the light cycle

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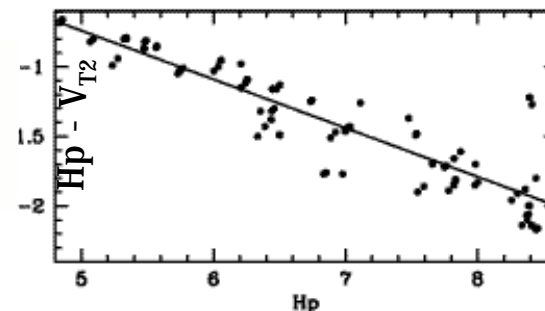
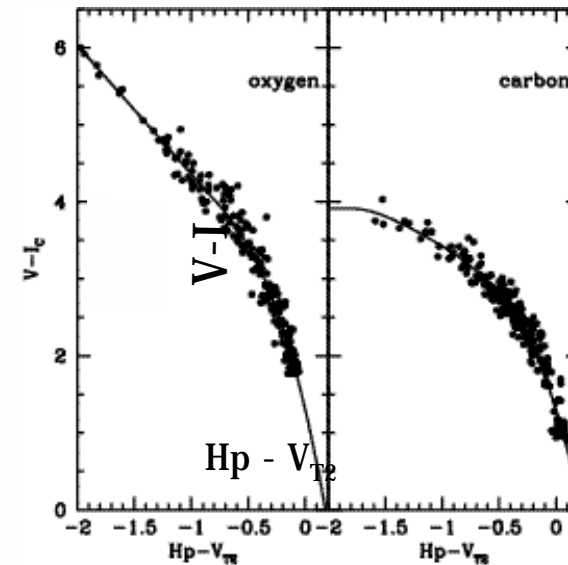
Example: Oxygen-rich Mira T Cep (HIP 104451)

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$$5.0 \leq H_p \leq 8.5$$



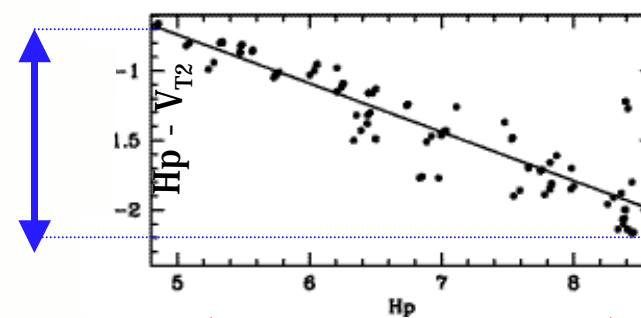
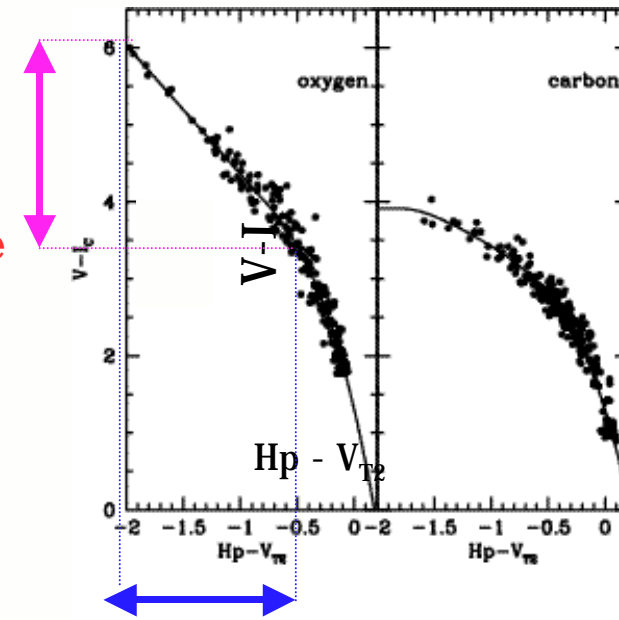
←→ Platais et al., 2003, A&A 397, 997

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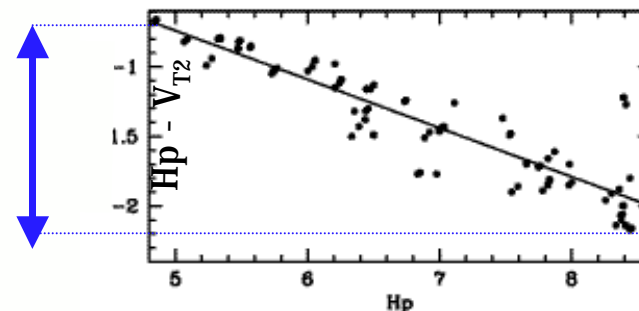
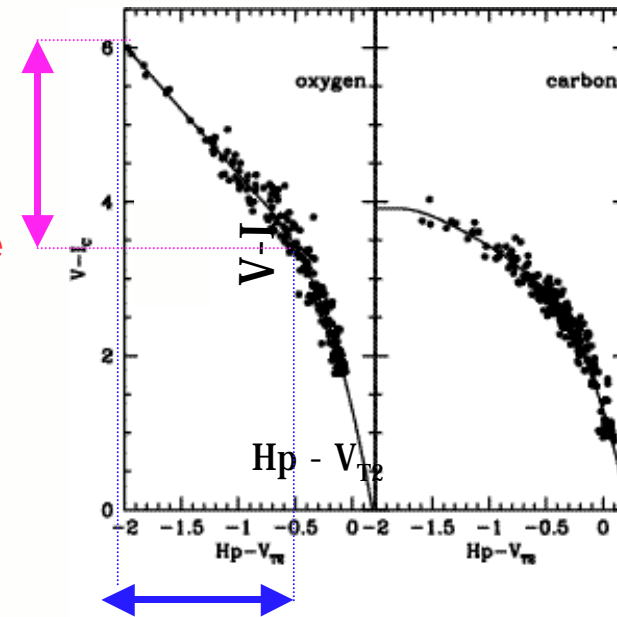
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$$5.0 \leq H_p \leq 8.5 \rightarrow 3 \leq V-I \leq 6$$



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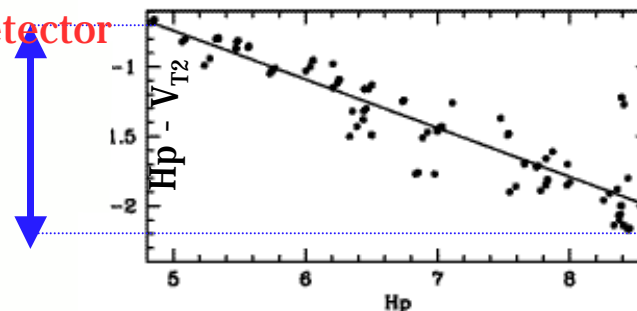
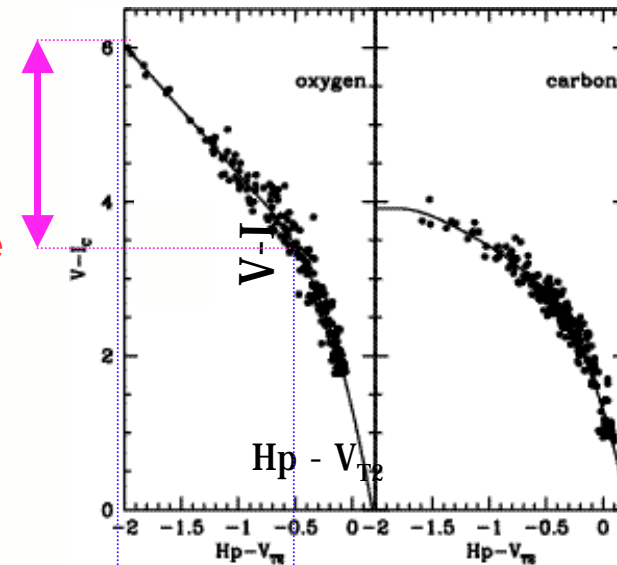
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Eqs. (16.8) and (16.12-16.18) of Vol. 3 of Hipparcos Catalogue
 photocenter moves by 1.9 mas on the Hipparcos detector



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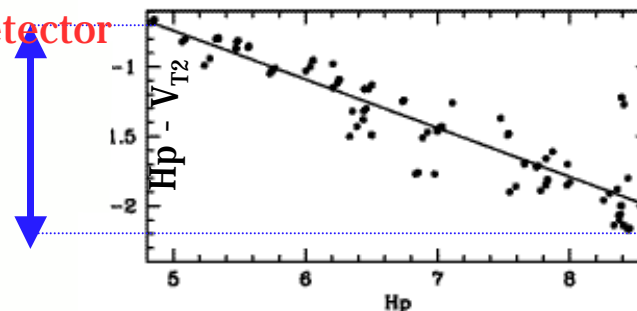
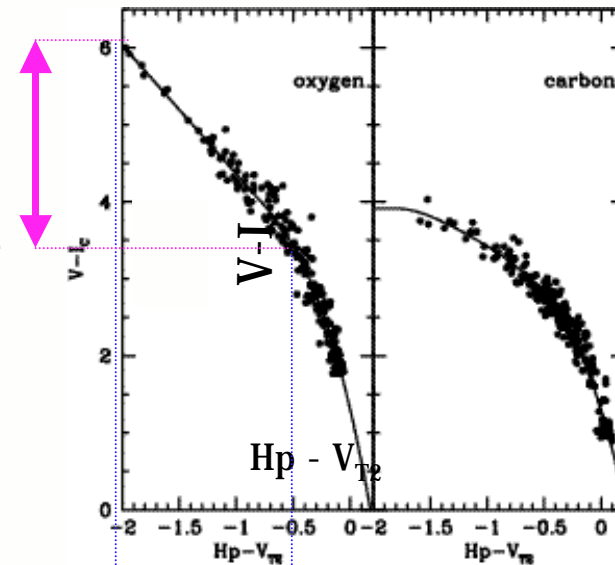
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Important that GAIA reduction pipeline uses
epoch photometry
 to perform the chromaticity correction



Astrometry of Long-Period Variable Stars: Solutions

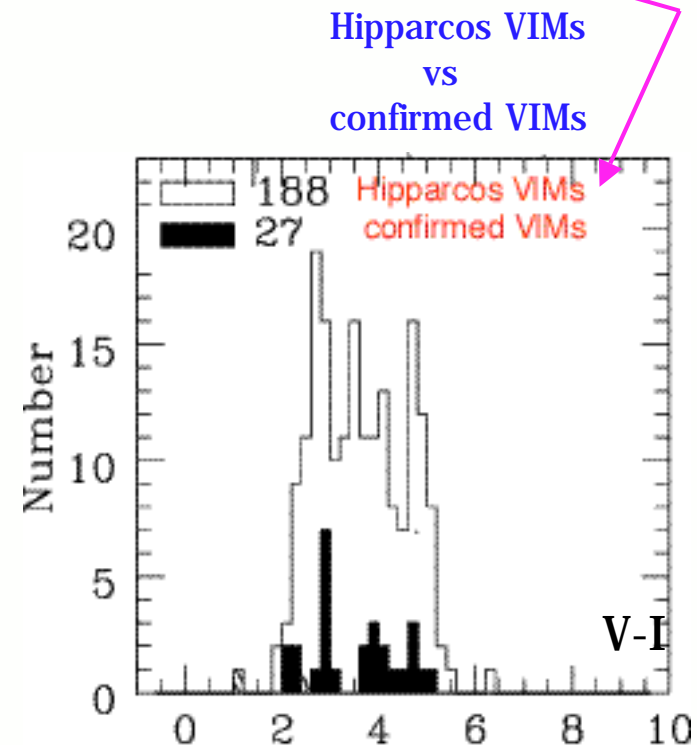
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When reduction pipeline uses *epoch V-I index* instead of average to perform the chromaticity correction, most *Variability-Induced Movers (VIMs)* disappear !!

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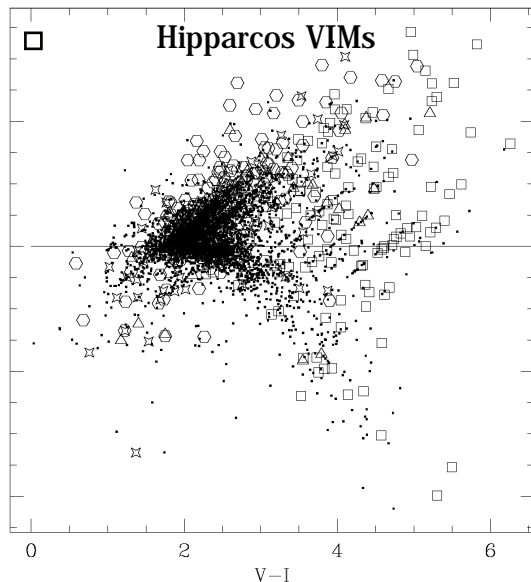


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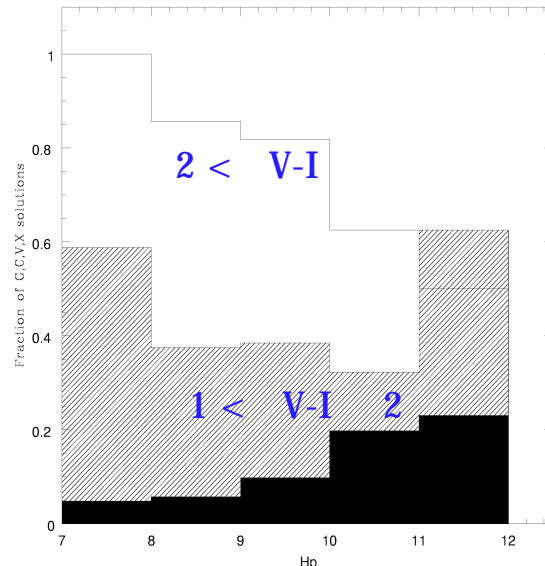
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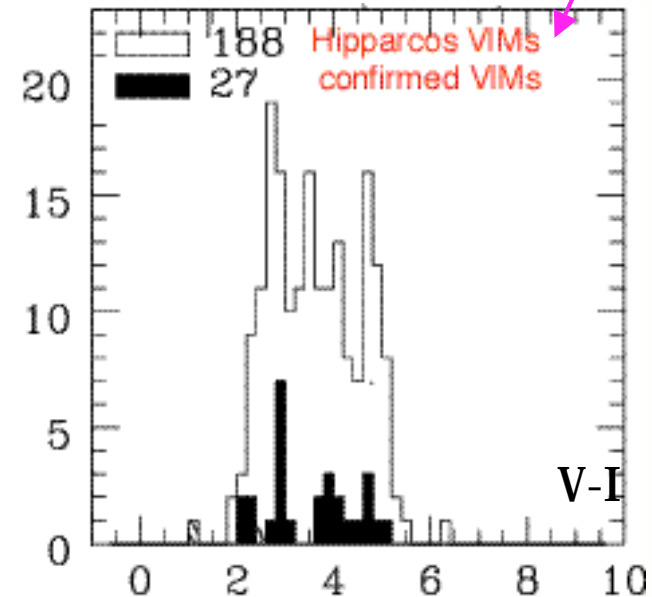
Hipparcos V-I (H75)
vs
Recalibrated V-I (from $H_p - V_{T2}$)



Fraction of Hipparcos VIMs vs H_p
for various V-I



Hipparcos VIMs
vs
confirmed VIMs



Astrometry of Long-Period Variable Stars: Solutions

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When reduction pipeline uses *epoch photometry* instead of average to perform the chromaticity correction, *anomalous parallaxes disappear !!*

Examples:

Oxygen-rich Mira R Hya (HIP 65835)

HIP = 1.62 ± 2.43 mas (VIM)
 revised = 8.44 ± 1.00 mas (non-VIM!)
 P-L = 7.14 mas

Pourbaix et al., 2003, A&A 399, 1167

Whitelock & Feast, 2000, MNRAS 319, 759

Oxygen-rich Mira T Cas (HIP 1834)

HIP = 0.59 ± 1.07 mas (VIM)
 revised = 3.08 ± 0.90 mas (non-VIM!)
 P-L = 3.45 mas
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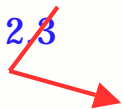
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The last word has not yet been said!