

Exercise 3: Vial Theorem

a) Derive the Reynolds transport theorem

[Mihalas & Mihalas p. 60]

Let I be the integral form of an equation, i.e.

$I = \int_{V(t)} \alpha d^3x$, where α is any single-valued scalar, vector, or tensor field. Also let $V(t)$ be some finite (material) volume with $dV(t=0) = dV_0$.

Furthermore, let us introduce the "color function" ψ :

$$\psi(\vec{x}) = \begin{cases} 1 & \vec{x} \in V(t) \\ 0 & \text{else} \end{cases}$$

such that
$$I = \int_{V(t)} \alpha d^3x = \int_{\mathbb{R}^3} \alpha \psi d^3x$$

This moves the t -dependence of the volume into the integral.

By definition,
$$\frac{D\psi}{Dt} = \frac{\partial \psi}{\partial t} + \vec{v} \cdot \nabla \psi = 0$$
 because ψ is moving along with the particle flow.

$$\begin{aligned}
 \text{Then: } \frac{dI}{dt} &= \frac{d}{dt} \left[\int_{V(t)} \alpha d^3x \right] = \frac{d}{dt} \left[\int_{\mathbb{R}^3} \alpha \psi d^3x \right] \\
 &= \int_{\mathbb{R}^3} \frac{d}{dt} [\alpha \psi] d^3x = \\
 &= \int_{\mathbb{R}^3} \left[\frac{\alpha}{t} \psi + \alpha \frac{\psi}{t} \right] d^3x
 \end{aligned}$$

Now use that $\frac{D\psi}{Dt} = 0 = \frac{\partial\psi}{\partial t} + \vec{v} \cdot \vec{\nabla}\psi \Rightarrow \frac{\partial\psi}{\partial t} = -\vec{v} \cdot \vec{\nabla}\psi$

$$\hookrightarrow \frac{dI}{dt} = \int_{\mathbb{R}^3} \left[\frac{\partial\alpha}{\partial t} \psi + \alpha \frac{\partial\psi}{\partial t} \right] d^3x = \int_{\mathbb{R}^3} \left[\frac{\partial\alpha}{\partial t} \psi - \alpha \vec{v} \cdot \vec{\nabla}\psi \right] d^3x$$

Furthermore use that:

$$\begin{aligned}
 \vec{\nabla} \cdot (\psi \alpha \vec{v}) &= \psi \vec{\nabla} \cdot (\alpha \vec{v}) + \alpha \vec{v} \cdot (\vec{\nabla}\psi) \\
 \Rightarrow -\alpha \vec{v} \cdot (\vec{\nabla}\psi) &= \psi \vec{\nabla} \cdot (\alpha \vec{v}) - \vec{\nabla} \cdot (\psi \alpha \vec{v})
 \end{aligned}$$

$$\begin{aligned}
 \hookrightarrow \frac{dI}{dt} &= \int_{\mathbb{R}^3} \left[\frac{\partial\alpha}{\partial t} \psi - \alpha \vec{v} \cdot (\vec{\nabla}\psi) \right] d^3x = \\
 &= \int_{\mathbb{R}^3} \left[\frac{\partial\alpha}{\partial t} \psi + \psi \vec{\nabla} \cdot (\alpha \vec{v}) - \vec{\nabla} \cdot (\psi \alpha \vec{v}) \right] d^3x \\
 &= \int_{\mathbb{R}^3} \left[\psi \left(\frac{\partial\alpha}{\partial t} + \vec{\nabla} \cdot (\alpha \vec{v}) \right) - \vec{\nabla} \cdot (\psi \alpha \vec{v}) \right] d^3x
 \end{aligned}$$

In the next step, show that $\int_{\mathbb{R}^3} \vec{\nabla} \cdot (\psi \alpha \vec{v}) d^3x = 0$ using the divergence theorem:

$$\int_V \vec{\nabla} \cdot \vec{f} d^3x = \int_{S(V)} \vec{f} \cdot \vec{n} dS$$

$$\int_{\mathbb{R}^3} \vec{\nabla} \cdot (\alpha \psi \vec{v}) d^3x = \int_{S(\mathbb{R}^3)} \alpha \psi \vec{v} \cdot \vec{n} d^3x$$

Remember that $\psi(\vec{x}) = 0$ for $x \notin V(t)$:

$$\Rightarrow \int_{S(\mathbb{R}^3)} \alpha \psi \vec{v} \cdot \vec{n} d^3x = 0$$

The surface that is being integrated over is clearly outside of $V(t)$.

$$\begin{aligned} \Rightarrow \frac{dI}{dt} &= \int_{\mathbb{R}^3} \psi \left[\frac{\partial \alpha}{\partial t} + \vec{\nabla} \cdot (\alpha \cdot \vec{v}) \right] d^3x \\ &= \int_{V(t)} \left[\frac{\partial \alpha}{\partial t} + \vec{\nabla} \cdot (\alpha \cdot \vec{v}) \right] d^3x \end{aligned}$$

This is the Reynolds transport theorem.

Now consider a specific variable, i.e. β for which is $\alpha = \rho \cdot \beta$ where ρ is the density.

$$\begin{aligned} \text{Then: } I(t) &= \int_{V(t)} \rho \beta d^3x \\ \dot{I}(t) &= \int_{V(t)} \left[\frac{\partial}{\partial t} (\rho \beta) + \vec{\nabla} \cdot (\rho \beta \vec{v}) \right] d^3x \\ &= \int_{V(t)} \left[\beta \frac{\partial \rho}{\partial t} + \rho \frac{\partial \beta}{\partial t} + \beta \vec{\nabla} \cdot (\rho \vec{v}) + (\rho \vec{v}) \cdot (\vec{\nabla} \beta) \right] d^3x \\ &= \int_{V(t)} \left[\rho \left(\frac{\partial \beta}{\partial t} + \vec{v} \cdot (\vec{\nabla} \beta) \right) + \beta \underbrace{\left(\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{v}) \right)}_{[*]} \right] d^3x \end{aligned}$$

If the Euler equations are used, then the term $[*]$ is the mass conservation equation: $\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{v}) = 0$

$$\Rightarrow \dot{I}(t) = \int_V \rho \left(\frac{\partial \beta}{\partial t} + \vec{v} \cdot (\vec{\nabla} \beta) \right) d^3x = \int_V \rho \frac{D\beta}{Dt} d^3x$$

b) Derive the virial theorem

Start with the "scalar moment of inertia": $I = \int_{V(t)} \rho |\mathbf{x}|^2 d^3x$

$$\begin{aligned} \text{Then: } \dot{I}(t) &= \int_{V(t)} \rho \frac{D}{Dt} (|\mathbf{x}|^2) d^3x = \int \rho \frac{D}{Dt} (\vec{x} \cdot \vec{x}) d^3x \\ &= \int_{V(t)} \rho 2 \cdot \vec{x} \cdot \frac{D\vec{x}}{Dt} d^3x = 2 \int \rho \vec{x} \cdot \vec{v} d^3x \end{aligned}$$

where it was used that $\frac{D\vec{x}}{Dt} = \vec{v}$
(by definition of the comoving derivation; see Mikhalas 55-56)

$$\Rightarrow \frac{1}{2} \dot{I} = \int_{V(t)} \rho \vec{x} \cdot \vec{v} d^3x$$

Then:

$$\begin{aligned} \frac{1}{2} \ddot{I} &= \frac{1}{2} \frac{d}{dt} \dot{I} = \int_{V(t)} \rho \frac{D(\vec{x} \cdot \vec{v})}{Dt} d^3x = \\ &= \int_{V(t)} \rho \left[\vec{v} \frac{D\vec{x}}{Dt} + \vec{x} \frac{D\vec{v}}{Dt} \right] d^3x = \int_{V(t)} \rho \left[v^2 + \vec{x} \frac{D\vec{v}}{Dt} \right] d^3x \end{aligned}$$

For the last part of the integral, use the Euler equations again: $\rho \frac{D\vec{v}}{Dt} = -\vec{\nabla} P + \rho \vec{a}$

$$\begin{aligned} \Rightarrow \int_{V(t)} \rho \vec{x} \frac{D\vec{v}}{Dt} d^3x &= \int_{V(t)} \vec{x} [-\vec{\nabla} P + \rho \vec{a}] d^3x = \int_{V(t)} \rho \cdot \vec{x} \cdot \vec{a} d^3x \\ &= \int_{V(t)} \rho \cdot \vec{x} \cdot \vec{a} d^3x - \int_{V(t)} \vec{x} \cdot \vec{\nabla} P d^3x \end{aligned}$$

For the last term, once again use that

$$\begin{aligned} \vec{\nabla} (P\vec{x}) &= \vec{x} (\vec{\nabla} P) + P (\underbrace{\vec{\nabla} \cdot \vec{x}}_{=3}) = \vec{x} (\vec{\nabla} P) + 3P \\ \rightarrow -\vec{x} \cdot (\vec{\nabla} P) &= 3P - \vec{\nabla} \cdot (P\vec{x}) \end{aligned}$$

Then:

$$\begin{aligned} - \int_{V(t)} \vec{x} \cdot \vec{\nabla} P d^3x &= \int_V 3P d^3x - \int_V \nabla \cdot (P\vec{x}) d^3x \quad [\text{divergence theorem again}] \\ &= \int_V 3P d^3x - \int_{S(V)} P\vec{x} \cdot \vec{n} dS \end{aligned}$$

Putting all together:

$$\frac{1}{2} \ddot{I} = \int_{V(t)} \rho \cdot v^2 d^3x + \int_{V(t)} \rho \cdot \vec{x} \cdot \vec{a} d^3x + \int_{V(t)} 3P d^3x - \int_{S(V)} P \vec{x} \cdot \vec{n} dS$$

$$= \underbrace{\int_V \rho \cdot v^2 d^3x}_{2 \cdot \text{Macroscopic kinetic energy}} + \underbrace{\int_V 3P d^3x}_{2 \cdot \text{Microscopic kinetic energy}} + \underbrace{\int_V \rho \cdot \vec{x} \cdot \vec{a} d^3x}_{\text{Force term: } \vec{a} \text{ comes from any force(s)}} - \underbrace{\int_S P \vec{x} \cdot \vec{n} dS}_{\text{Surface term}}$$

2 · Macroscopic kinetic energy
 $2 \cdot \frac{1}{2} \rho v^2 dV$

2 · Microscopic kinetic energy
 $2 \cdot \frac{3}{2} P dV$
" $\equiv 2 \cdot \frac{3}{2} N k_B dT$ "

Force term:
 \vec{a} comes from any force(s)

Surface term

← my "interpretation"

$$\underbrace{\hspace{15em}}_{\equiv 2K}$$

$$\underbrace{\hspace{15em}}_{\equiv V}$$

$$\underbrace{\hspace{15em}}_{\equiv S}$$

$$\frac{1}{2} \ddot{I} = 2K + V - S$$

For an equilibrium state, we must demand: $\ddot{I} = 0$

[Note: $\dot{I} = 2 \int_V \vec{x} \cdot \rho \cdot \vec{v} d^3x$ represents the sum of ^{the dot products of} all particle's positions and their velocities, so it sort of represents the system's total action. Demanding $\ddot{I} = 0$ would mean that the total action of the system is constant over time.

Furthermore: A system in equilibrium must be bound. If it is, then

$$\langle \dot{I} \rangle = \frac{1}{T} \int_0^T \dot{I} dt = \frac{1}{T} (I(T) - I(0)) \xrightarrow{T \rightarrow \infty} 0$$

The average only diverges if $I(t)$ is not bound ($I(T) \rightarrow \infty$), which would mean $\vec{v} \rightarrow \infty$ since $I = \int_V \rho |\vec{x}|^2 d^3x$

c) Stellar mass - radius relation

Assuming equilibrium state and a uniform, spherical system

1) Calculate gravitational acceleration

$$\text{Poisson equation: } \Delta \varphi = 4\pi G \rho = \vec{\nabla} \cdot (\vec{\nabla} \varphi) = \vec{\nabla} \cdot (-\vec{g}) = -\vec{\nabla} \cdot \vec{g}$$

$$\text{Assume spherical symmetry: } \vec{g} = g \vec{e}_r$$

Integrate Poisson equation:

$$\begin{aligned} \int_V \Delta \varphi dV &= \int_V -\vec{\nabla} \cdot \vec{g} dV \stackrel{\text{div. th.}}{=} - \int_S \vec{g} \cdot \vec{n} dS = - \int g \cdot \vec{e}_r \cdot \vec{e}_r dS = - \int g dS = -4\pi r^2 g \\ &= \int_V 4\pi G \rho dV = \frac{16\pi^2 G \rho r^3}{3} \end{aligned}$$

$$\Rightarrow g = \frac{-4\pi G \rho r^3}{3r^2} = -\frac{4\pi G \rho r}{3}$$

2) Virial theorem: $\frac{1}{2} \ddot{I} = 2K + V - S$

$$\text{Assuming equilibrium: } \ddot{I} = 0$$

$$\text{Assuming static system: } \vec{v} = 0$$

$$K = \int_V \frac{1}{2} \rho v^2 d^3x + \underbrace{\int_V \rho v^2 d^3x}_=0 = \frac{3}{2} \int_V \rho d^3x = \frac{3}{2} \frac{4\pi}{3} R^3 \rho = 2\pi R^3 \rho$$

$$\begin{aligned} V &= \int_V \rho (\vec{g} \cdot \vec{x}) d^3x = \int_V \rho \cdot g \cdot \vec{e}_r \cdot r \cdot \vec{e}_r \cdot r^2 \sin\theta dr d\theta d\phi = \\ &= -\frac{4\pi}{3} G \rho^2 \cdot 4\pi \int_0^R r^4 dr = -\frac{16\pi^2}{15} G \rho^2 R^5 \end{aligned}$$

$$S = \int_S p \vec{x} \cdot \vec{n} d^3x = 0 \quad \text{no pressure on surface}$$

$$\Rightarrow 2K + V = 0 = 4\pi R^3 \rho - \frac{16\pi^2}{15} G \rho^2 R^5$$

$$\Rightarrow \rho = \frac{4\pi}{15} G \rho^2 R^2$$

3) Chose an equation of state

E.g. isothermal EOS: $T = T_0$

$$PV = N k_B T_0$$

With $\frac{N}{V} = \frac{\rho}{m}$ [or $N = \frac{M}{m}$; $N/V = \frac{M/V}{m} = \frac{\rho}{m}$]

$$P = \frac{\rho}{m} k_B T_0$$

$$= \frac{4\pi}{15} G \rho^2 R^2 \quad (\text{from before})$$

$$\Rightarrow \frac{k_B T_0}{m} = \frac{4\pi}{15} G \rho R^2 =$$

$$= \frac{4\pi}{15} G R^2 \cdot \frac{M}{\frac{4\pi R^3}{3}} = \underline{\underline{\frac{1}{5} G \frac{M}{R}}}$$

mass - radius relation

Note: Other EOS can be chosen. The isothermal one is the simplest. The point is that it needs a relation between the pressure P and the density, so the total mass can be calculated.