

## Dynamical Timescale

Characterizes departures from mechanical equilibrium, and free-fall. In free-fall, the internal pressure is negligible.

$$\ddot{r} = \frac{1}{8} \frac{\partial P}{\partial S} - \frac{GM_r}{r^2} = - \frac{GM_r}{r^2}$$

i) By approximations:

$$\ddot{r} \sim \frac{R_c - R_s}{\tau_{\text{dyn}}^2} = \frac{-R}{\tau_{\text{dyn}}^2} \sim - \frac{GM}{R^2}$$

$$\Rightarrow \tau_{\text{dyn}} \sim \sqrt{\frac{R^3}{GM}} \sim \sqrt{\frac{1}{G\rho}}$$

ii) More exact: again  $\frac{dP}{dr} \approx 0$

$$\ddot{r} = - \frac{GM_r}{r^2}$$

$$\Rightarrow \ddot{r} \dot{r} = - \frac{GM_r}{r^2} \dot{r}$$

$$\Rightarrow \frac{d}{dt} \left( \frac{1}{2} \dot{r}^2 \right) = \frac{d}{dt} \left( \frac{GM_r}{r} \right)$$

$$\Rightarrow \frac{1}{2} \dot{r}^2(t) - \dot{r}^2(t_{\text{ini}}) = GM_r \left( \frac{1}{r} - \frac{1}{r_{\text{ini}}} \right)$$

Assume  $t_{ini} = 0$ ,  $r(t_{ini}) = r_{ini} = \infty$ ,  $\dot{r}(t_{ini}) = 0$ ,

$$M = \frac{4}{3} \pi r^3 \bar{\rho}$$

$$\Rightarrow \frac{1}{2} \dot{r}^2(t) = \frac{G}{r} \frac{4}{3} \pi r^3 \bar{\rho}$$

$$\Rightarrow \dot{r}^2(t) = \frac{8\pi G \bar{\rho}}{3} r^2$$

$$\Rightarrow \frac{\dot{r}(t)}{r} = -\sqrt{\frac{8\pi G \bar{\rho}}{3}} \quad \text{take minus for contraction}$$

Assuming  $M_r = \text{const} = 4\pi \bar{\rho} r^3$ :

$$\frac{dM_r}{M_r} = 0 = \frac{4\pi r^3 d\bar{\rho}}{4\pi r^3 \bar{\rho}} + \frac{4\pi \bar{\rho} 3r^2 dr}{4\pi r^3 \bar{\rho}}$$

$$\Rightarrow \frac{d\bar{\rho}}{\bar{\rho}} = -3 \frac{dr}{r}$$

and

$$\frac{\dot{r}}{r} = \frac{1}{r} \frac{dr}{dt} = -\frac{1}{3} \frac{1}{\bar{\rho}} \frac{d\bar{\rho}}{dt} = -\sqrt{\frac{8\pi G \bar{\rho}}{3}}$$

$$\Rightarrow \frac{1}{\bar{\rho}} \frac{d\bar{\rho}}{dt} = \sqrt{24\pi G \bar{\rho}}$$

We define

$$\tau_{\#} = \bar{\rho} \frac{dt}{d\bar{\rho}} = \frac{1}{\sqrt{24\pi G \bar{\rho}}}$$