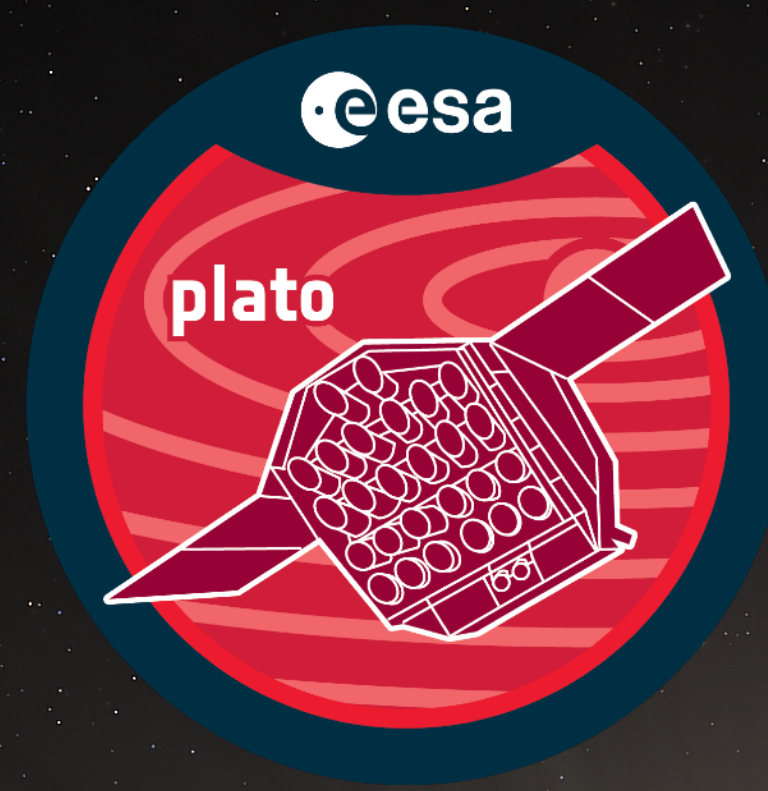




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DE GENÈVE**

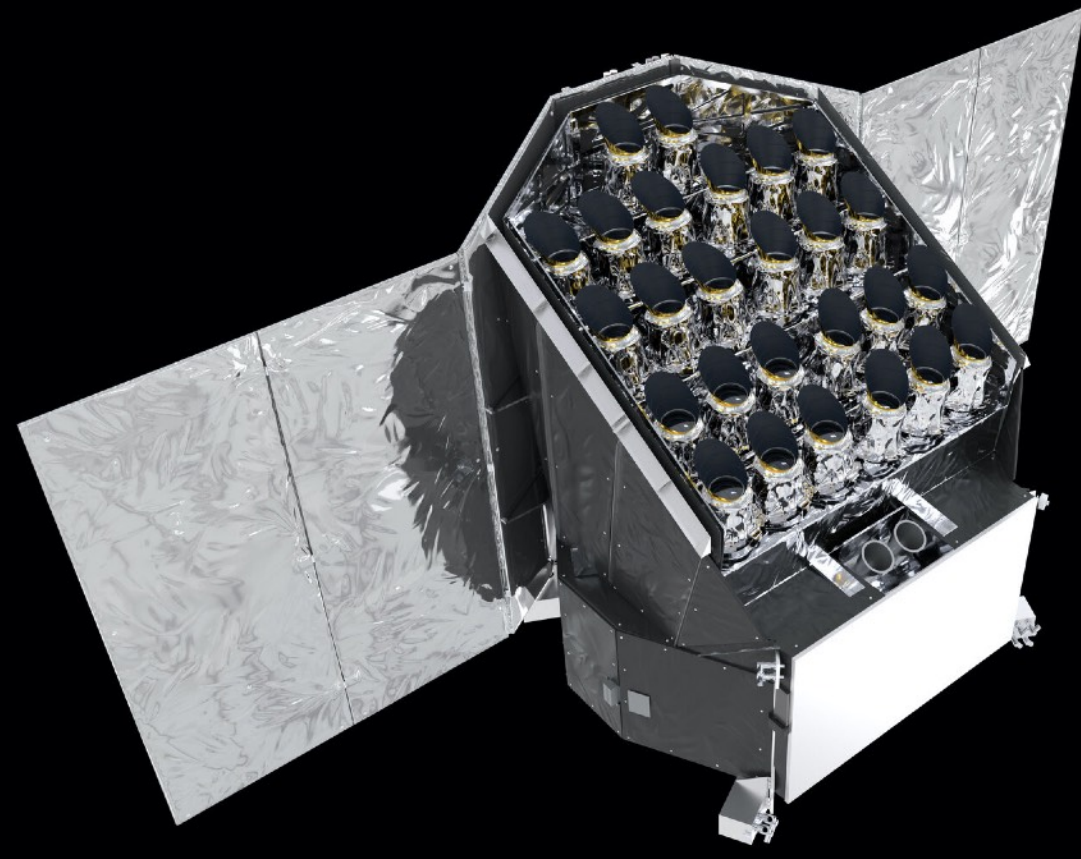


**PlanetS**

# Use of dynamical knowledge to improve follow-up efficiency

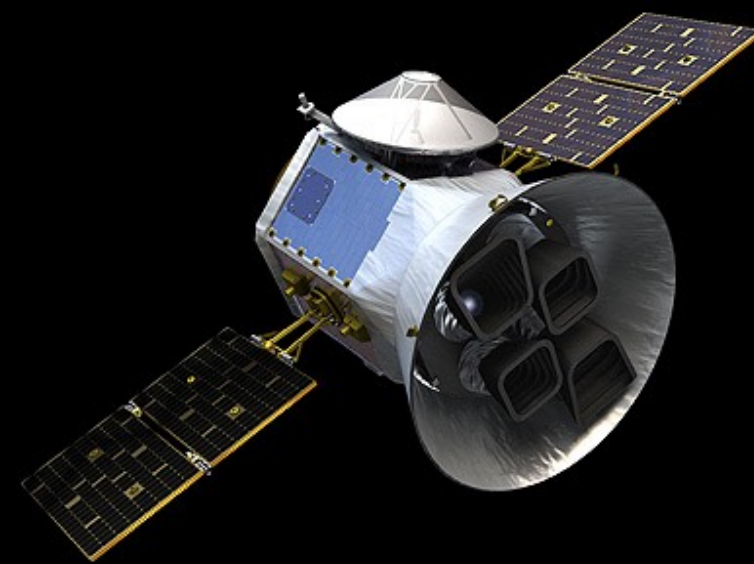
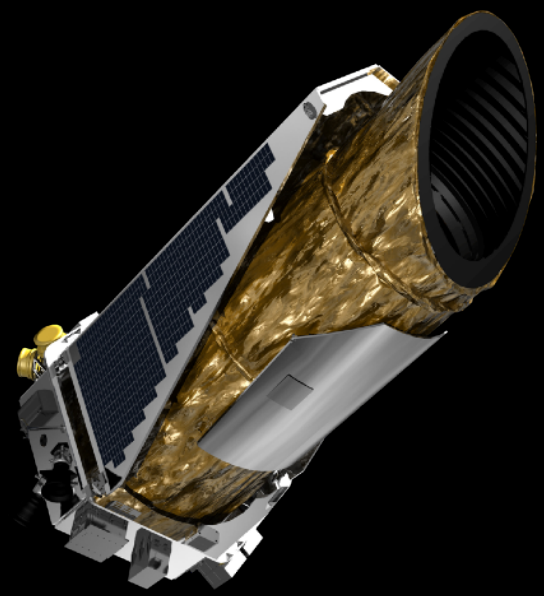
PLATO GOP Workshop 2022 - Adrien Leleu

# Cases



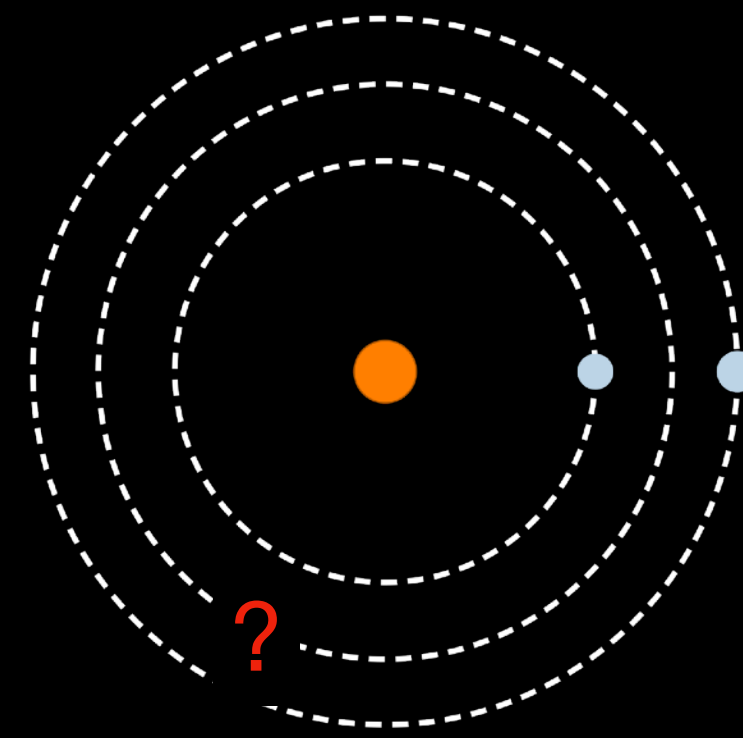
Continuous observations  
example: Kepler

Step and stare  
example: TESS

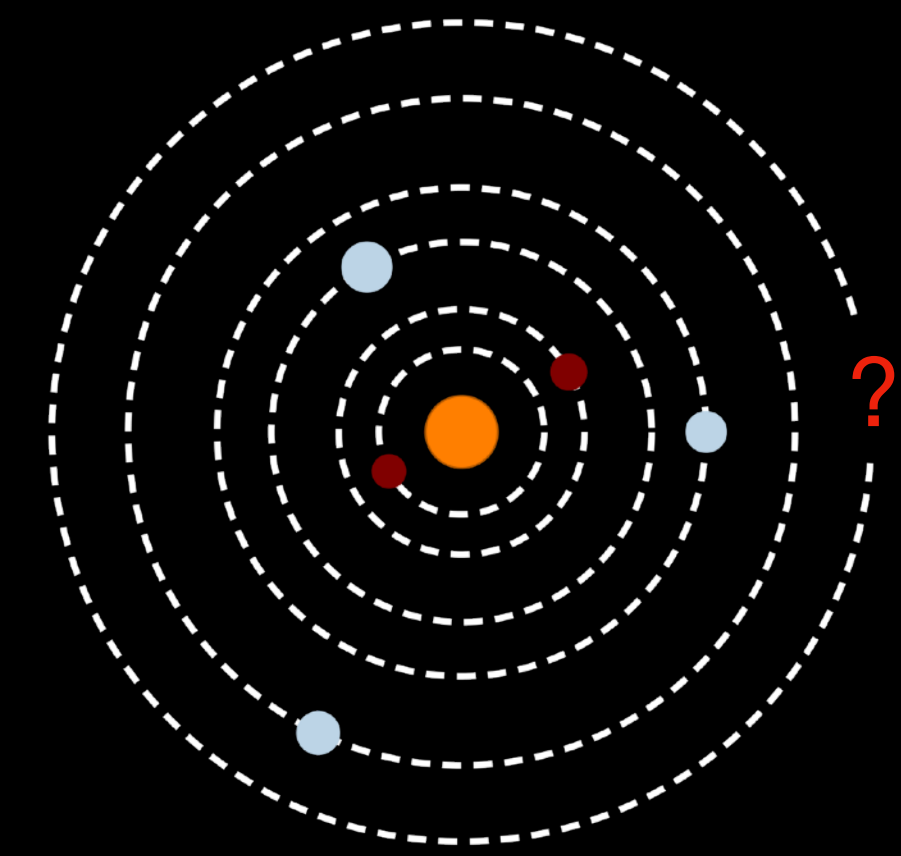


Only **periods** and **t0s** necessary,  
can be pushed to hundreds of days

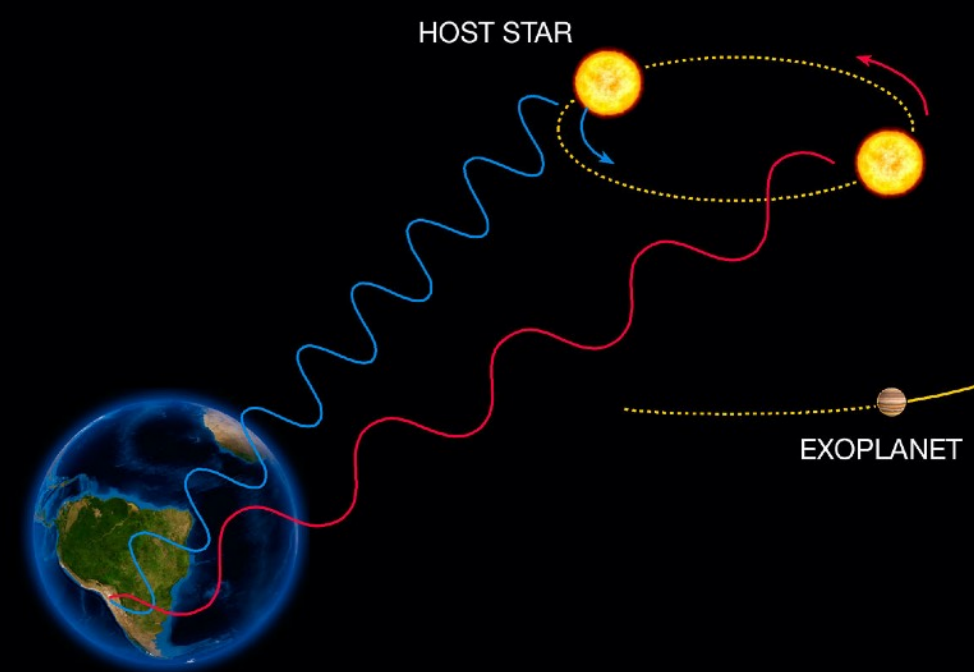
Verifying a  
potential resonant chain



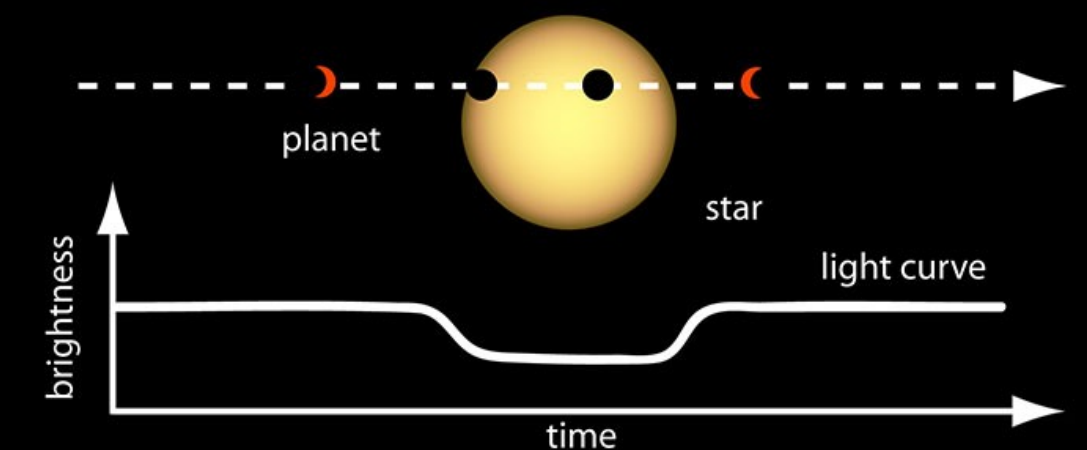
Completing a  
resonant system



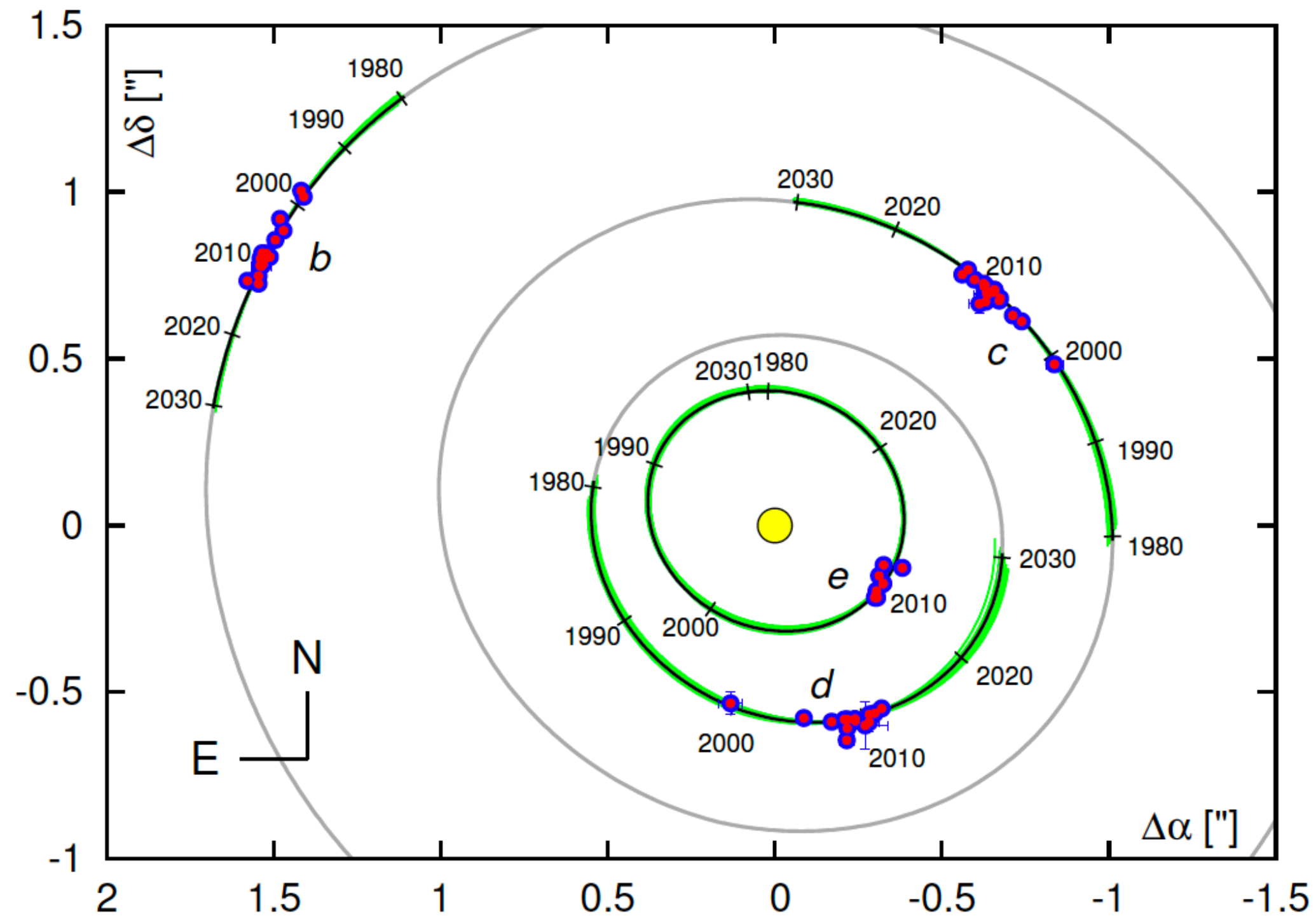
Radial velocity follow-up



Photometry follow-up



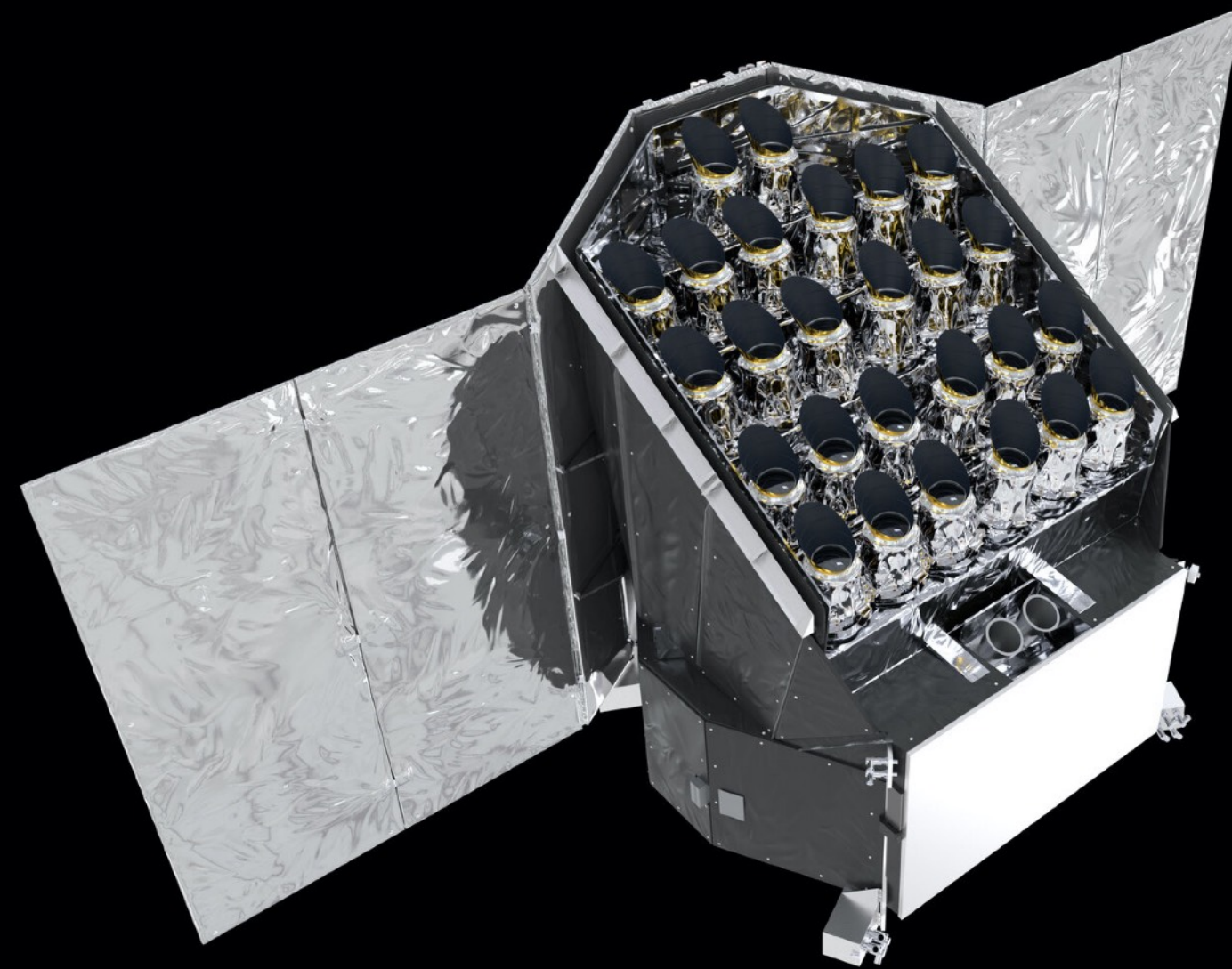
# Can take place at any period: HR 8799



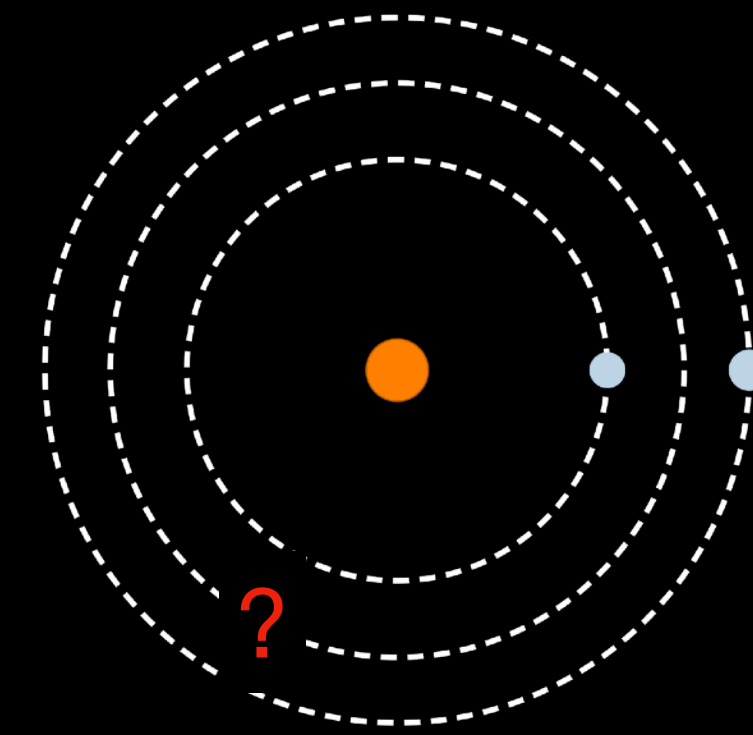
**Figure 4.** Relative astrometric positions of the planets (red filled circles), orbital arcs for the best-fitting model IVa (black curves), and stable solutions within the  $3\sigma$  confidence level of the best-fitting model (green curves).

	$m [m_{\text{Jup}}]$	$a [\text{au}]$
HR 8799 e	$9 \pm 2$	$15.4 \pm 0.2$
HR 8799 d	$9 \pm 3$	$25.4 \pm 0.3$
HR 8799 c	$9 \pm 3$	$39.4 \pm 0.3$
HR 8799 b	$7 \pm 2$	$69.1 \pm 0.2$

# Case 1



verifying a  
potential resonant chain



# Mean motion resonances (MMRs)

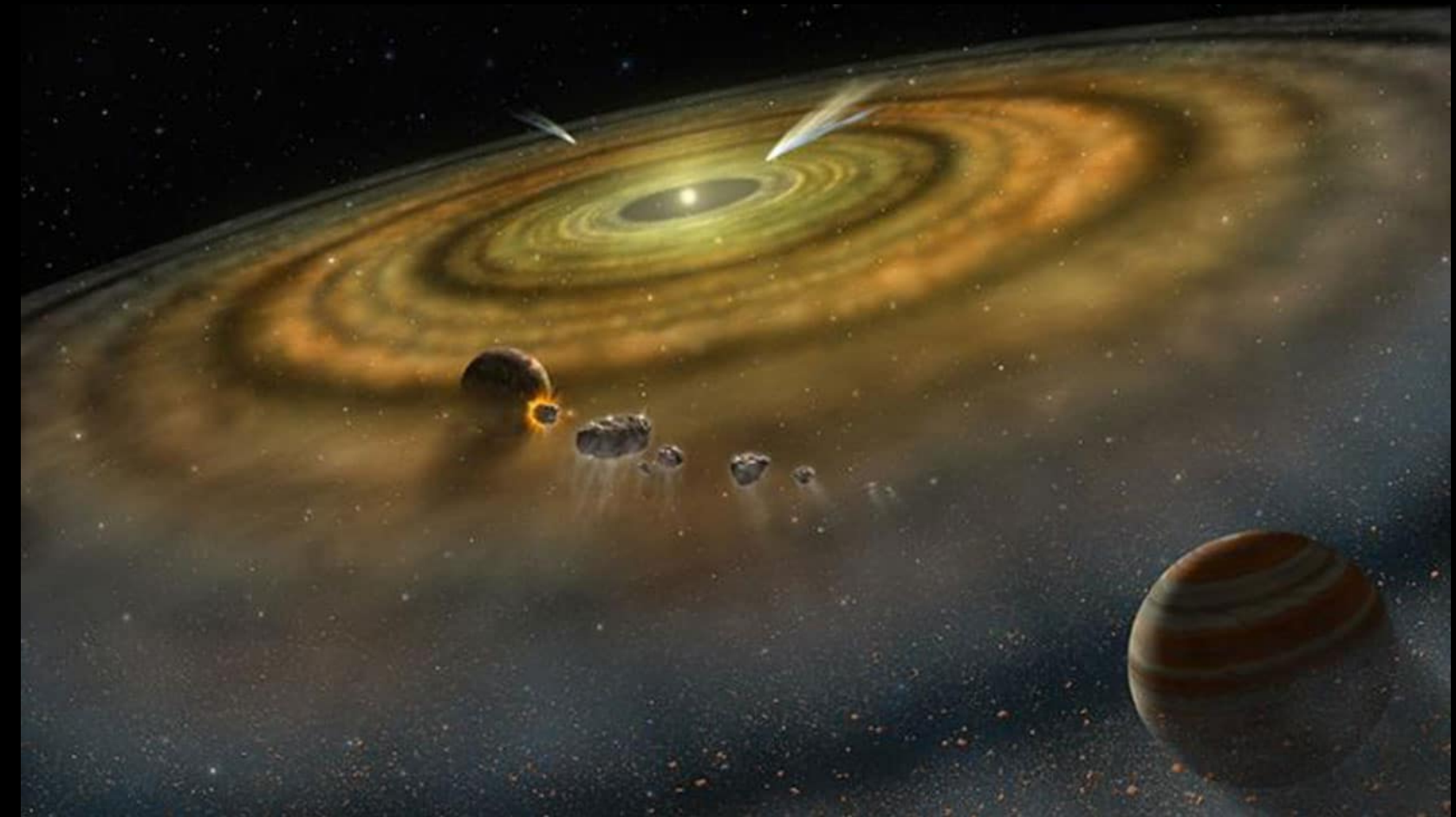
Byproduct of the formation of  
planetary systems

Two planets MMRs (first order):

$$P_2/P_1 \approx (k + 1)/k = 2, 3/2, 4/3, \text{ etc. } k \text{ integer}$$

Pluto - Neptune :

$$247.94 \text{ yr} / 164.8 \text{ yr} = 1.504 \approx 3/2$$



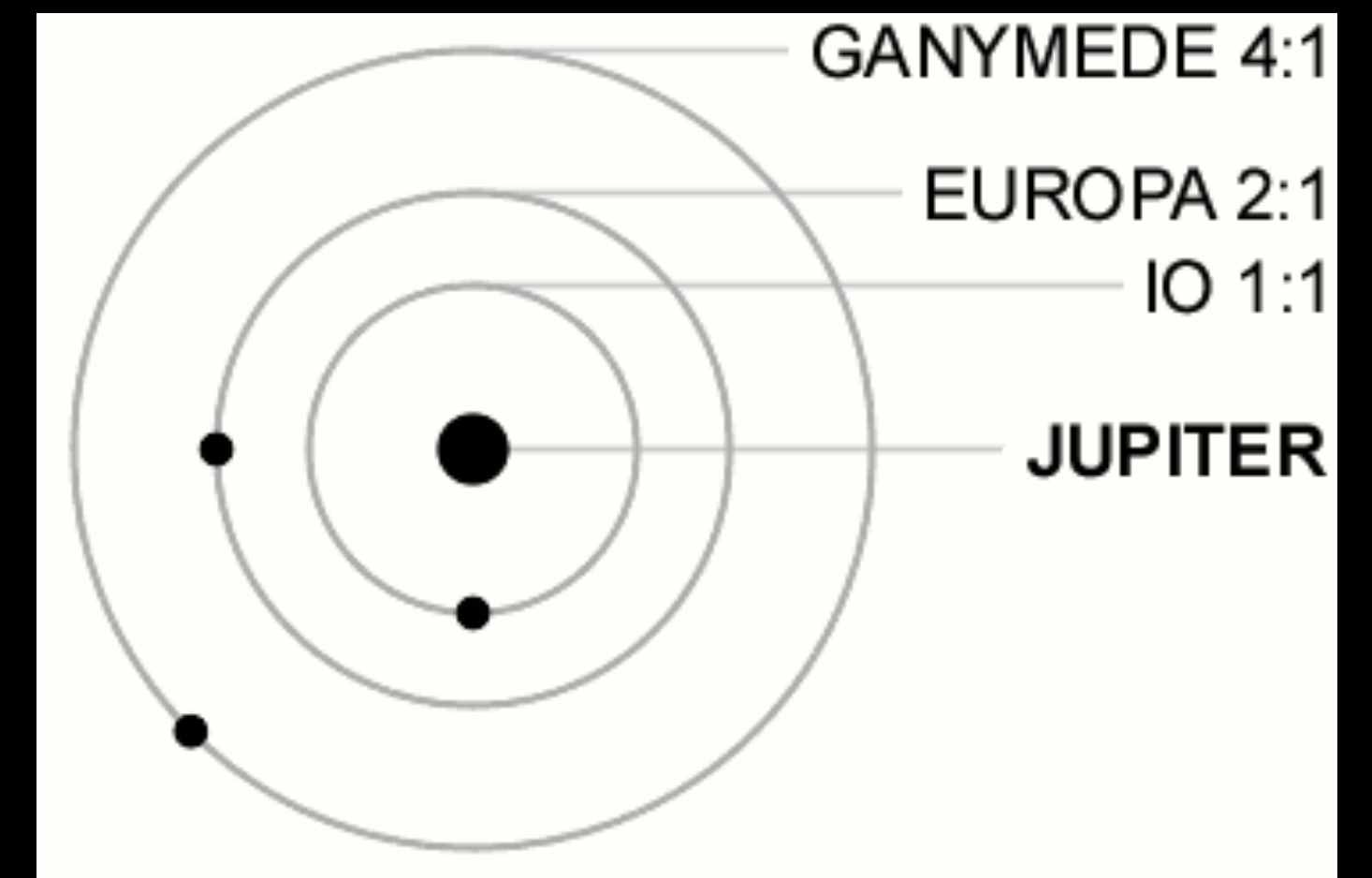
# Mean motion resonances (MMRs)

## (Generalised) Laplace resonance :

$$l/P_1 - (l + m)/P_2 + m/P_3 \approx 0 \quad \text{Laplace relation}$$

Io - Europa - Ganymede :

$$3/(3.52 \text{ hour}) - 1/(1.76 \text{ hour}) - 2/(7.15 \text{ hour}) \approx 0$$



## Chain of 2-planet MMRs:

$$k_1/P_1 - (k_1 + 1)/P_2 \approx 0$$

$$k_2/P_2 - (k_2 + 1)/P_3 \approx 0$$

$$\implies k_1/P_1 - (k_1 + k_2 + 1)/P_2 + (k_2 + 1)/P_3 \approx 0$$

Laplace relation

# Completion of a Laplace chain

First order MMRs are most commons:

$$P_2/P_1 \approx (k + 1)/k$$

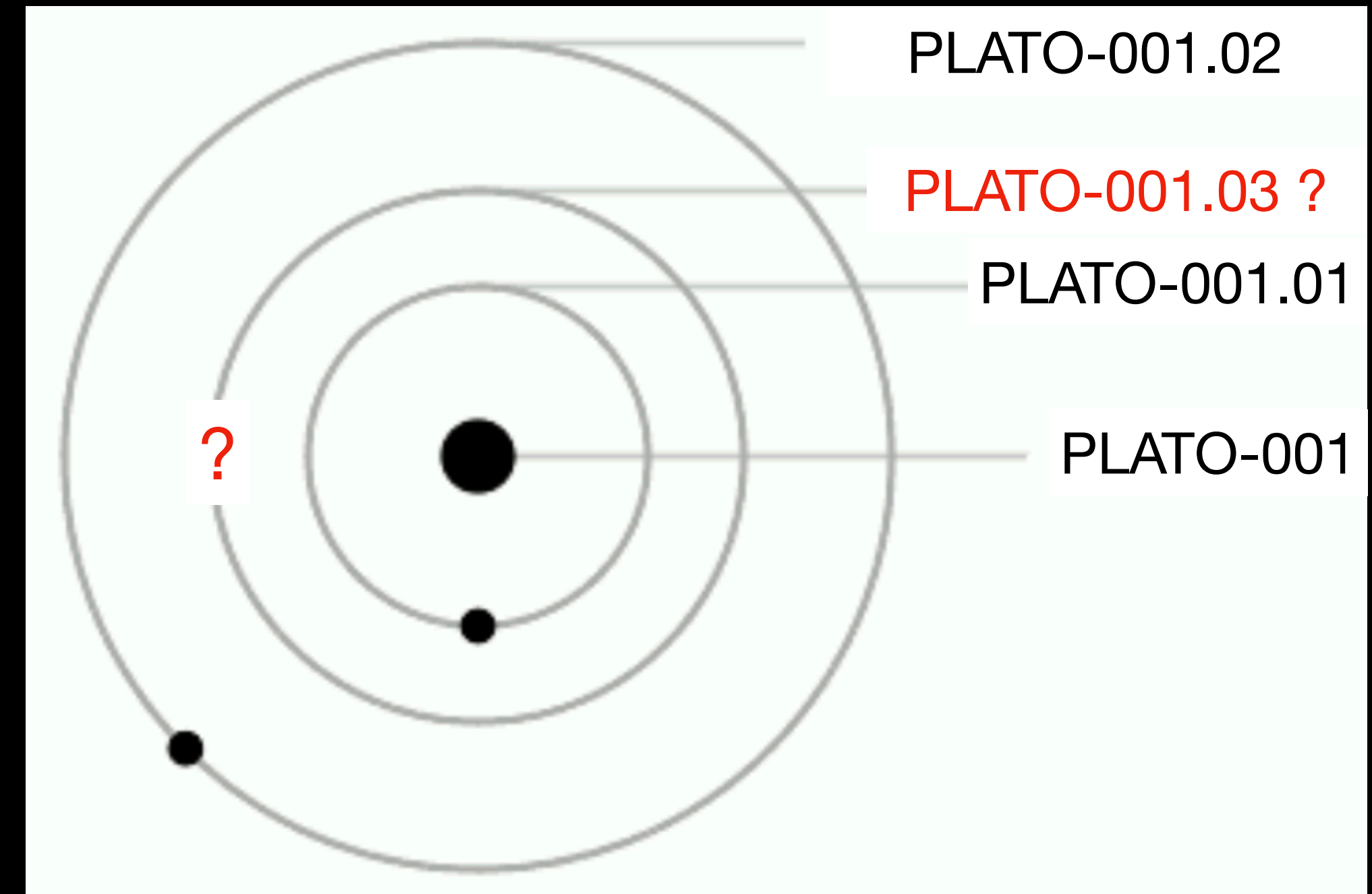
Second order MMRs are more rare:

$$P_2/P_1 \approx (k + 2)/k$$

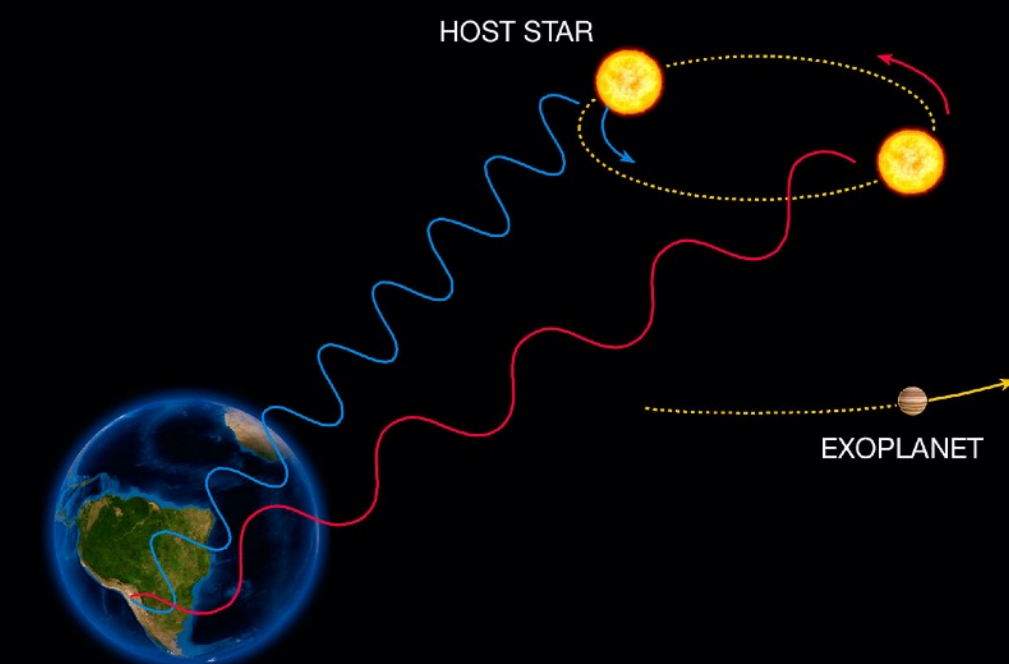
Higher order MMRs  $P_2/P_1 \approx (k + q)/k$ ,  $q \geq 3$  ex. 9/4, 4/1, 5/2

are very weak at low eccentricities.

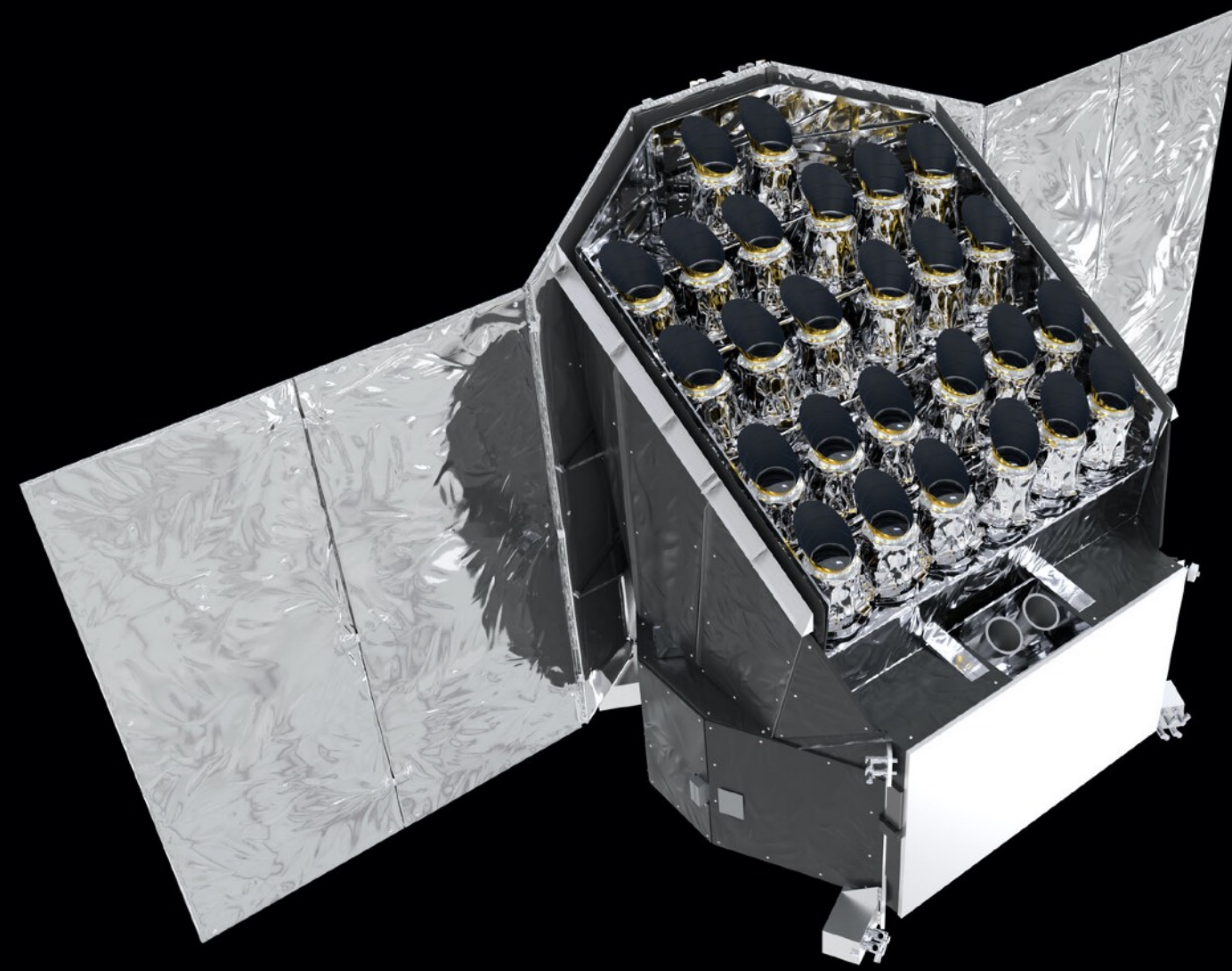
Example, if  $P_2/P_1 \approx 4.004$  if might hide and intermediate planet...



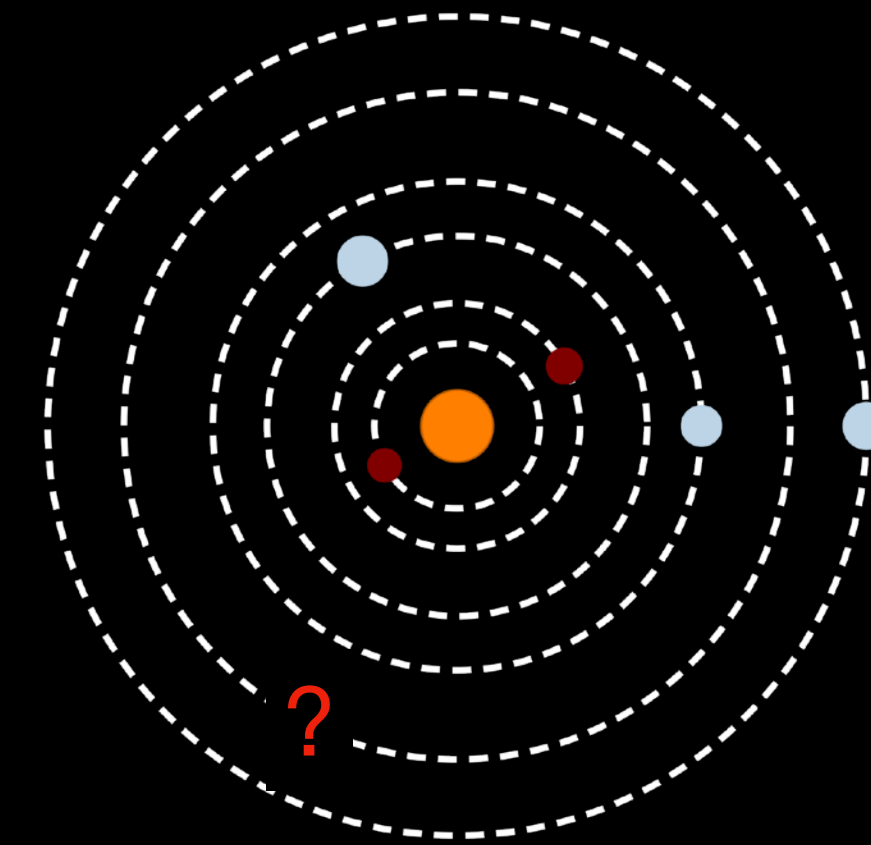
Radial velocity follow-up



# Case 2.1



Completing a  
resonant system

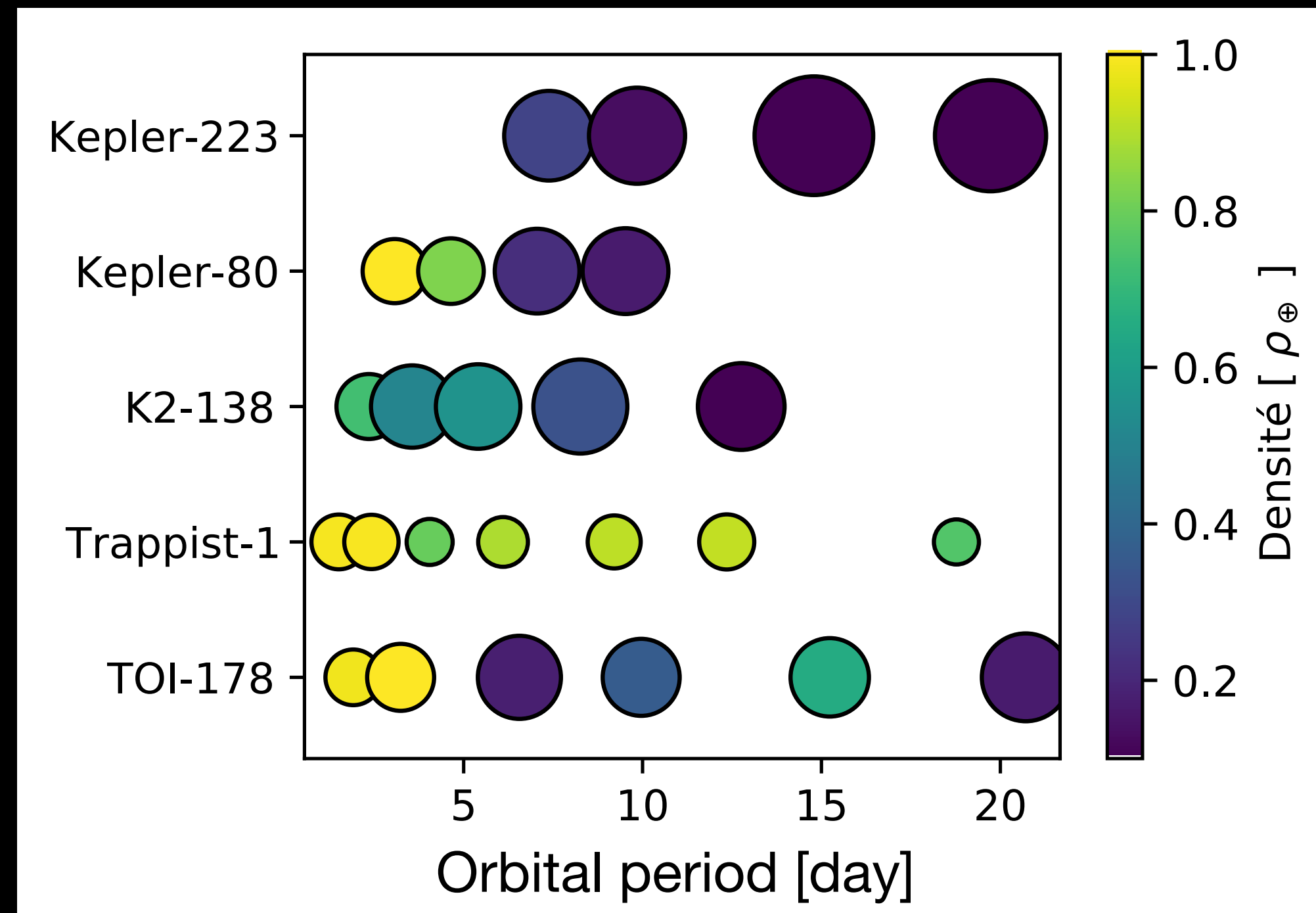


Planet inside the chain,  
single transit



# Chains of (generalised) Laplace resonances

4+ planets

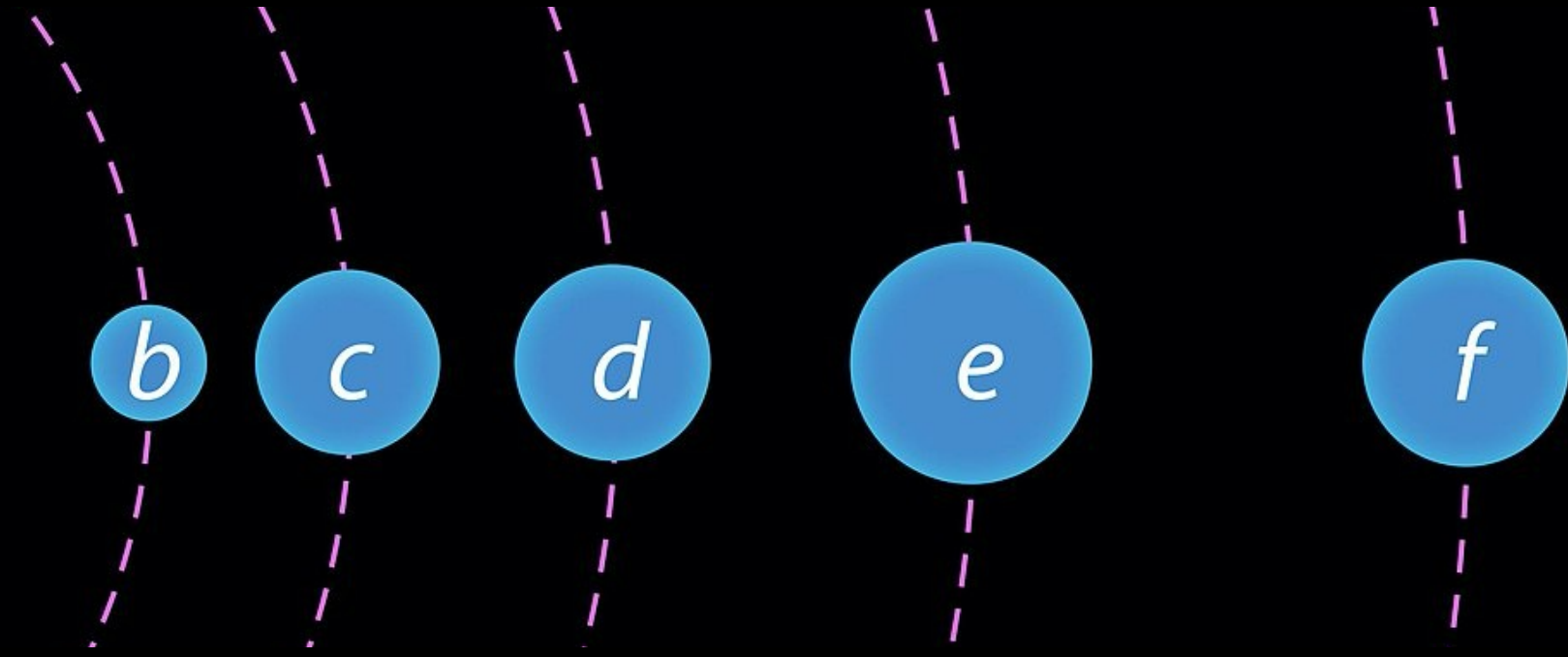


Use of the resonant chain to predict the period of a planet

Trappist-1 : prediction of Trappist-1 h (Luger+ 2017)  
TOI-178 : prediction of TOI-178 f (Leleu+ 2021)

TOI-1136 this morning (Dai+ 2022) !!

# K2-138 distance to 2-body MMRs



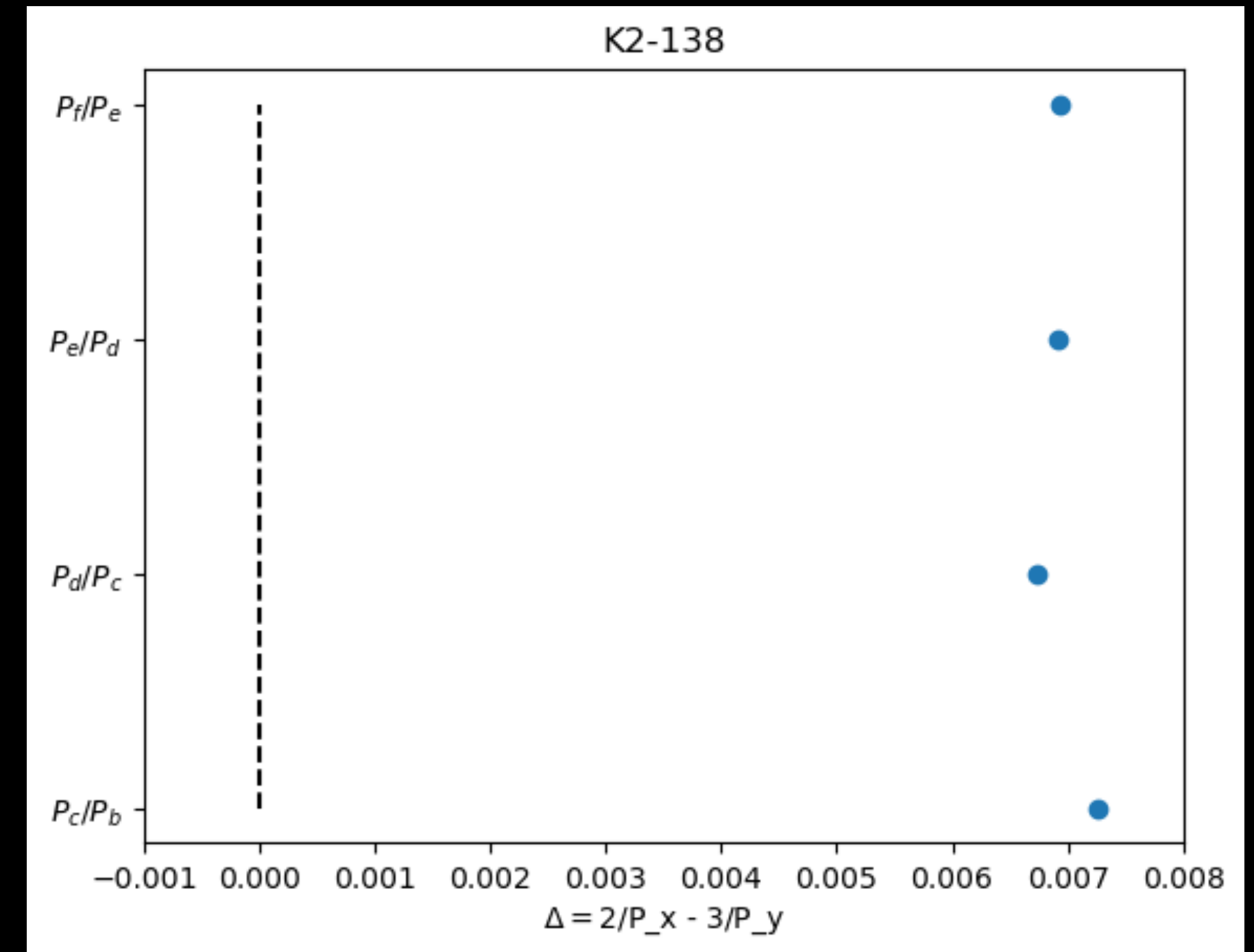
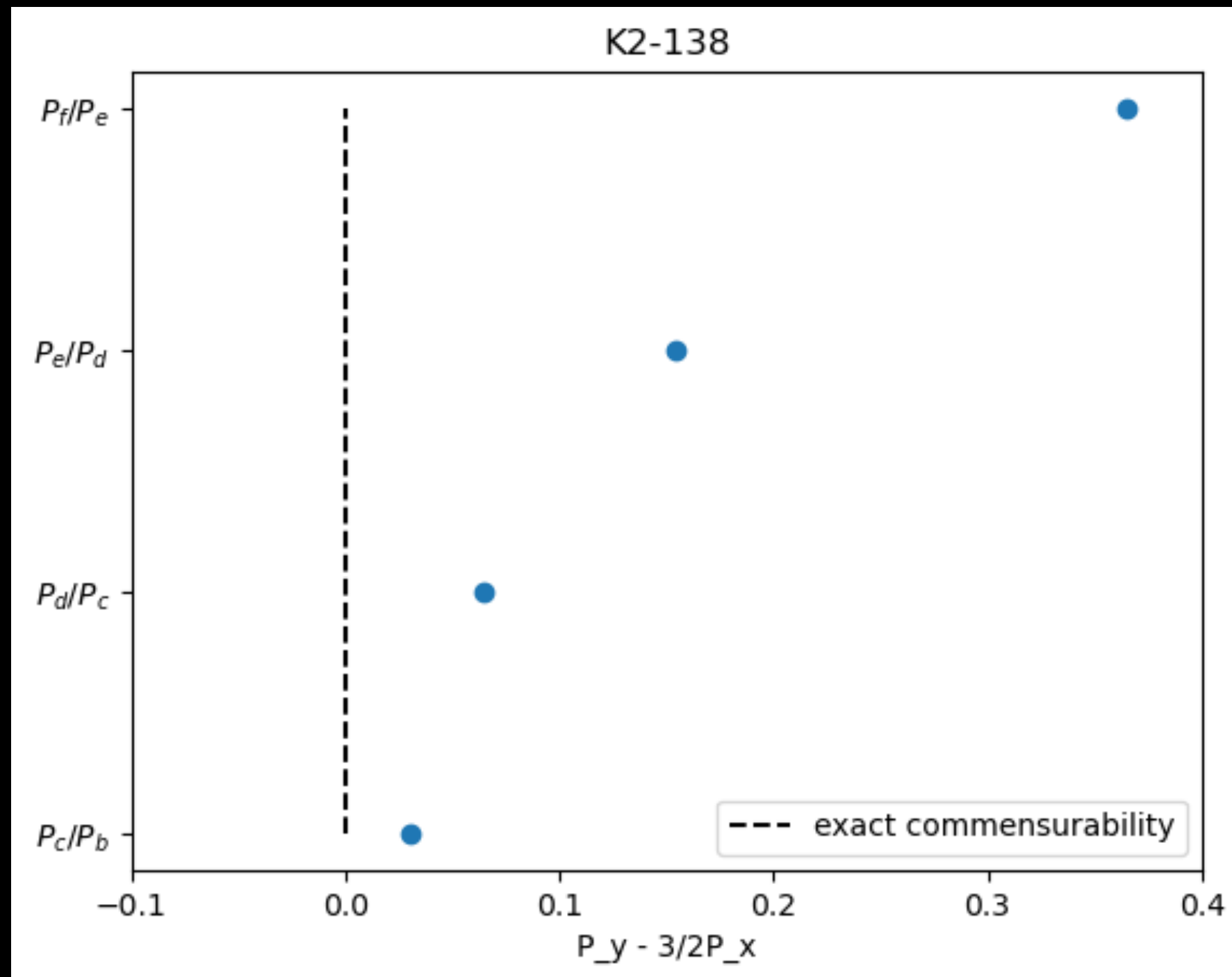
*Planet sizes enlarged  
50x relative to star*

(Lopez+ 2019)

$$\frac{P_c}{P_b} \approx 3/2 \quad \approx 3/2 \quad \approx 3/2 \quad \approx 3/2$$

Distance in period

Distance in frequency

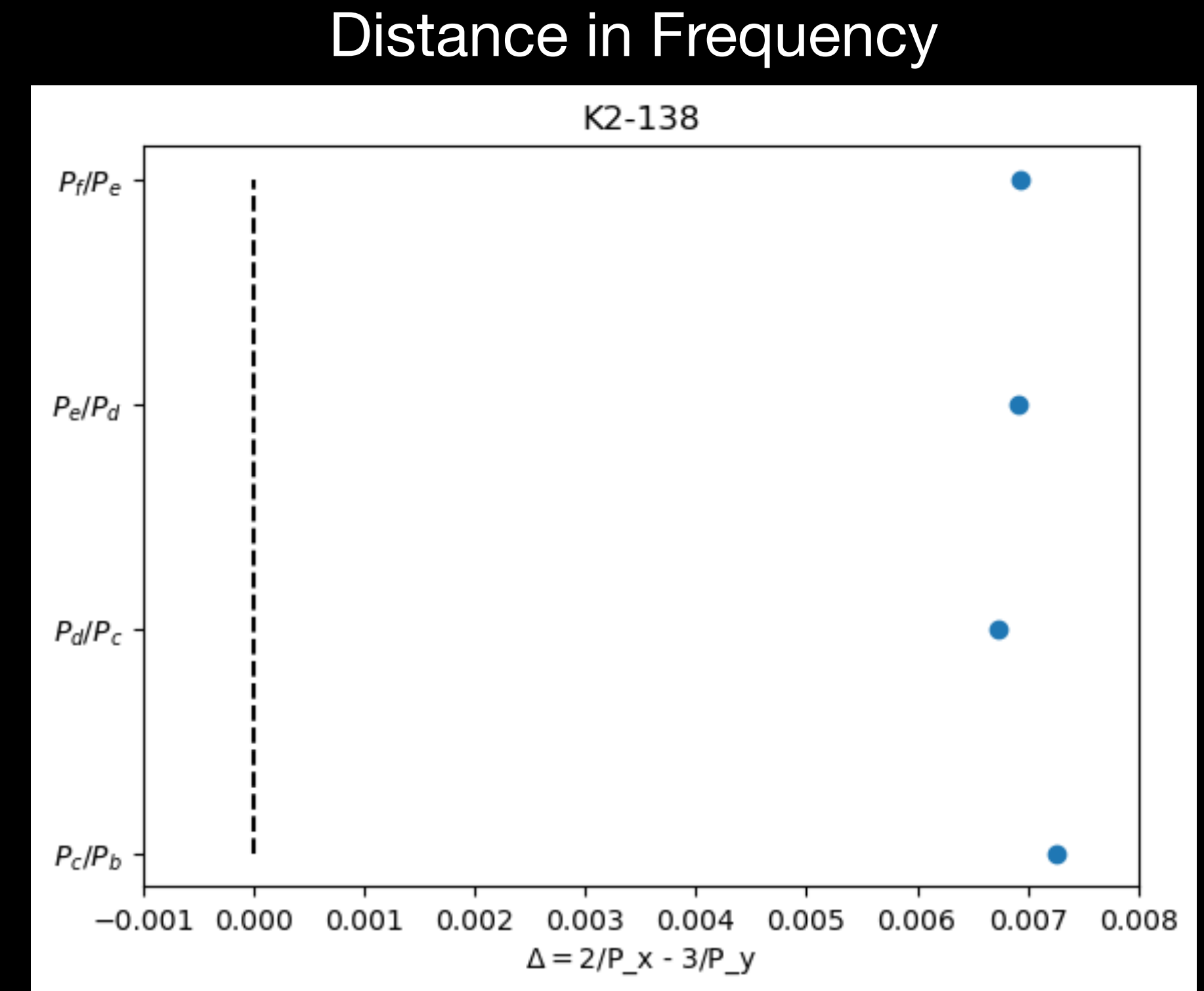


# Laplace resonances

Distance to the exact commensurability  
in frequency

$$k_1/P_1 - (k_1 + 1)/P_2 = \Delta_1$$

$$k_2/P_2 - (k_2 + 1)/P_3 = \Delta_2$$



$$\Delta_1 = \Delta_2 \implies k_1/P_1 - (k_1 + k_2 + 1)/P_2 + (K_2 + 1)/P_3 = 0$$

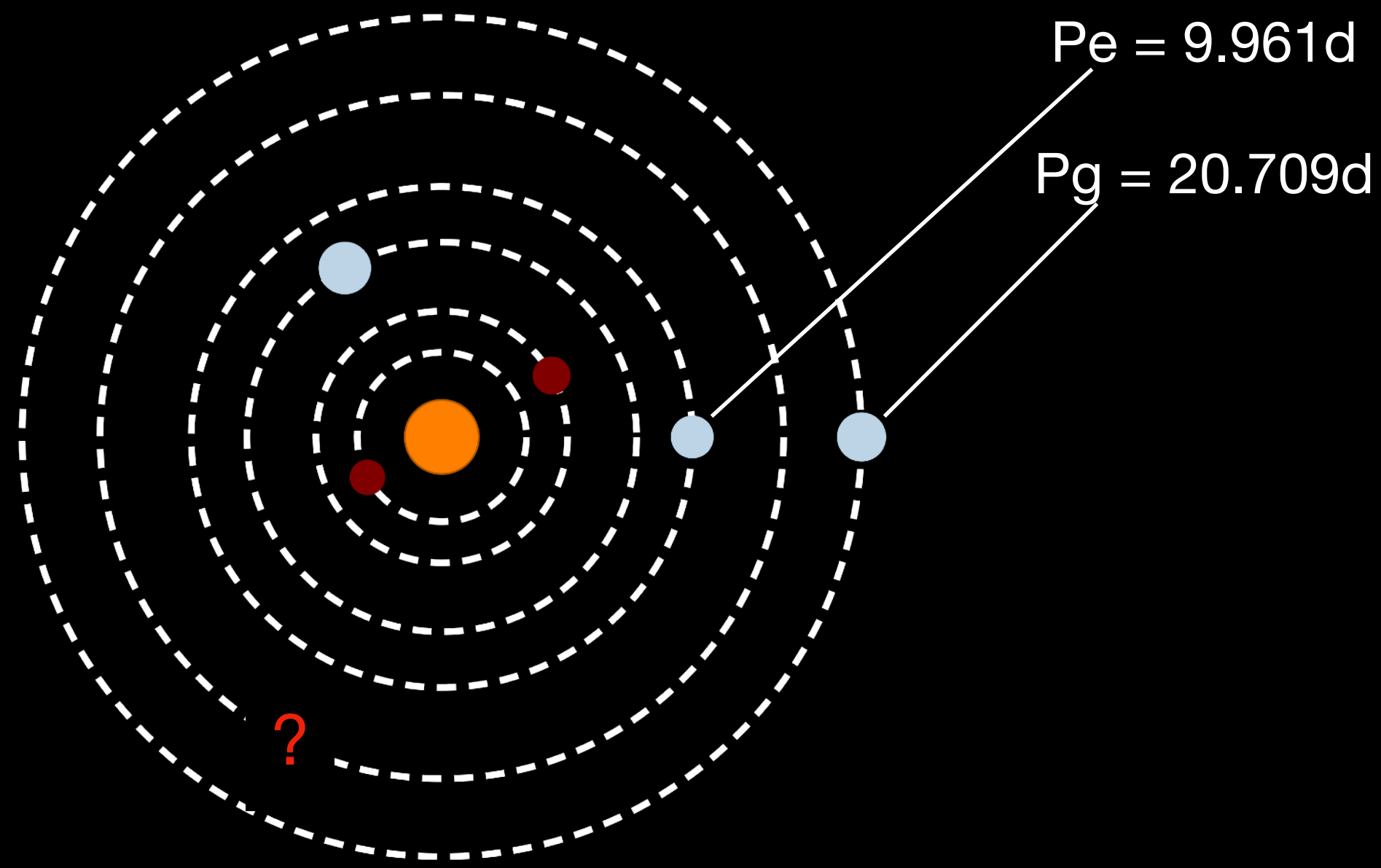
Laplace relation

Associated super period:  $P_{Super,12} = \frac{1}{\Delta_1} = \frac{1}{k_1/P_1 - (k_1 + 1)/P_2}$

(Lithwick+ 2012)

Have to be the same for all pairs in the chain

# prediction of the period of TOI-178f

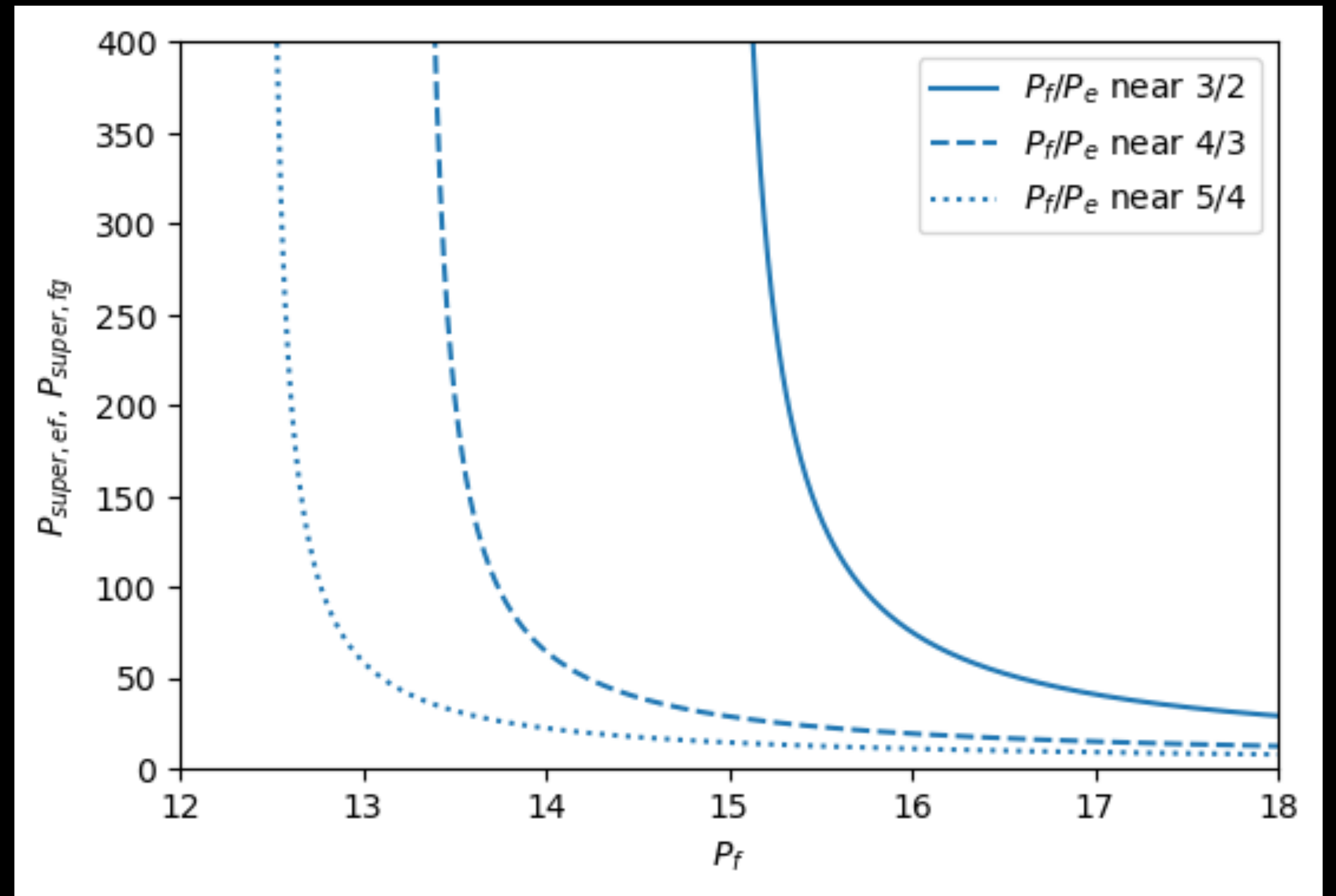


Laplace relation if:

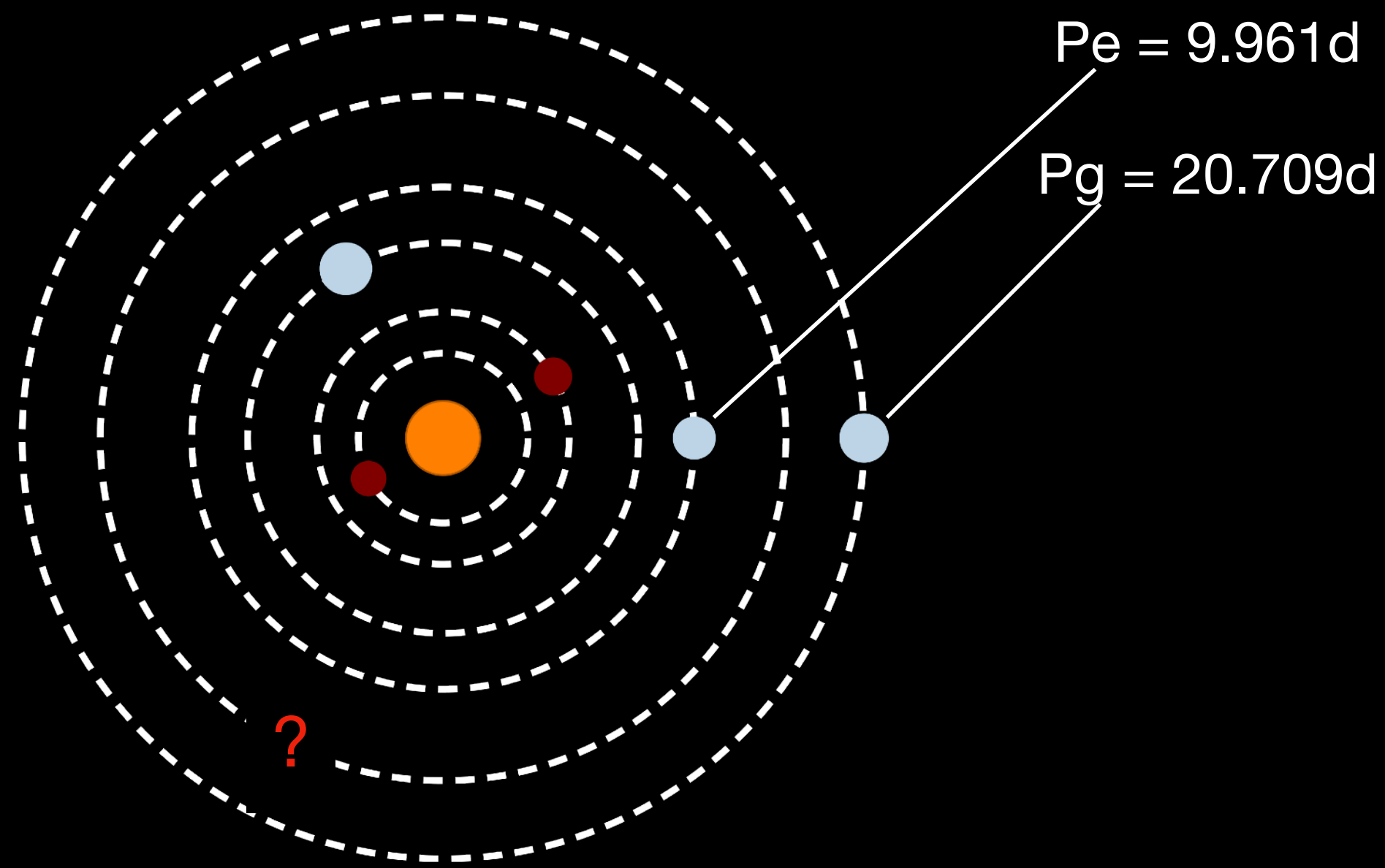
$$\Delta_1 = \Delta_2$$

$$P_{Super,ef} = P_{Super,fg}$$

$$P_{Super,12} = \frac{1}{\Delta_1} = \frac{1}{k_1/P_1 - (k_1 + 1)/P_2}$$



# prediction of the period of TOI-178f

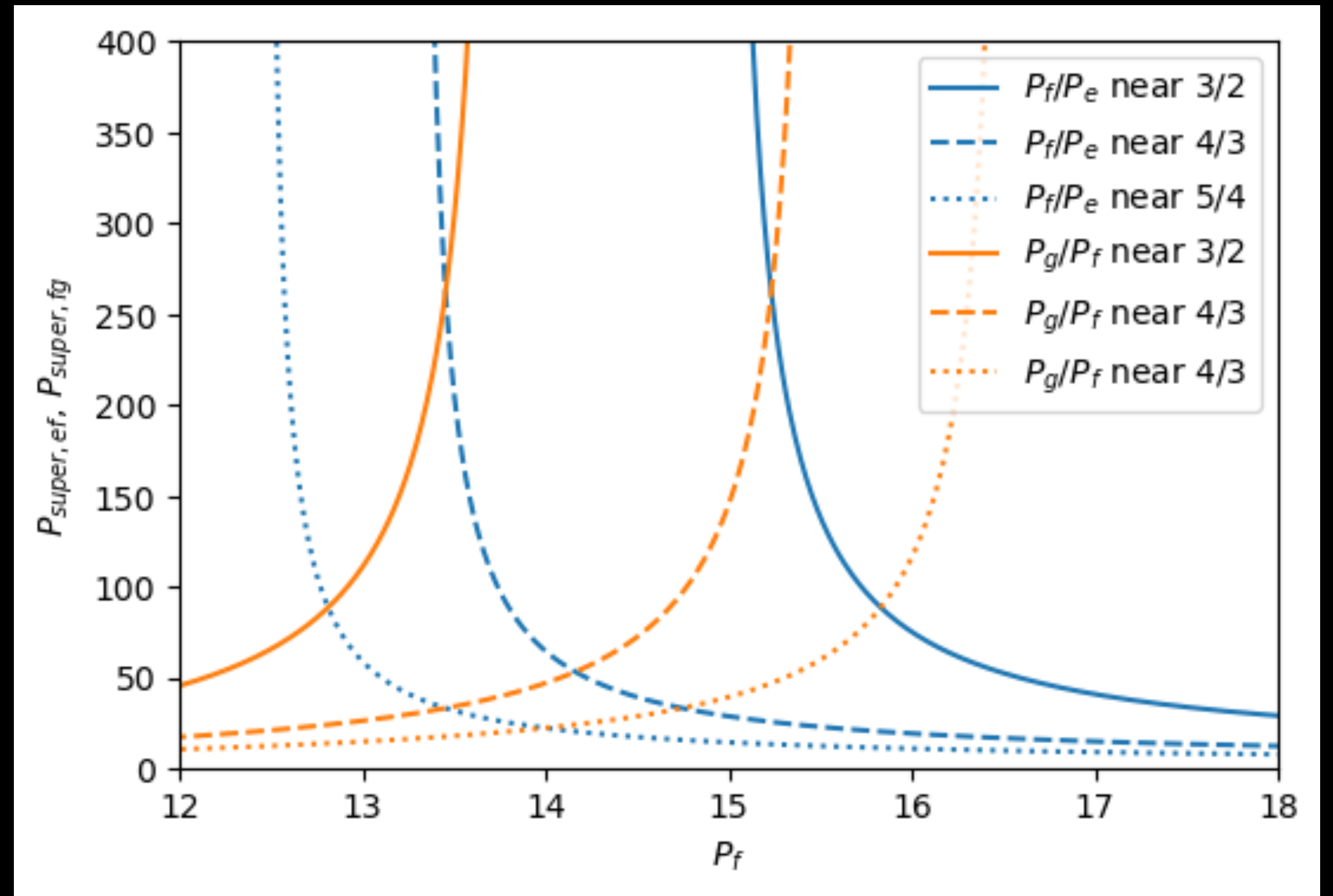


Laplace relation if:

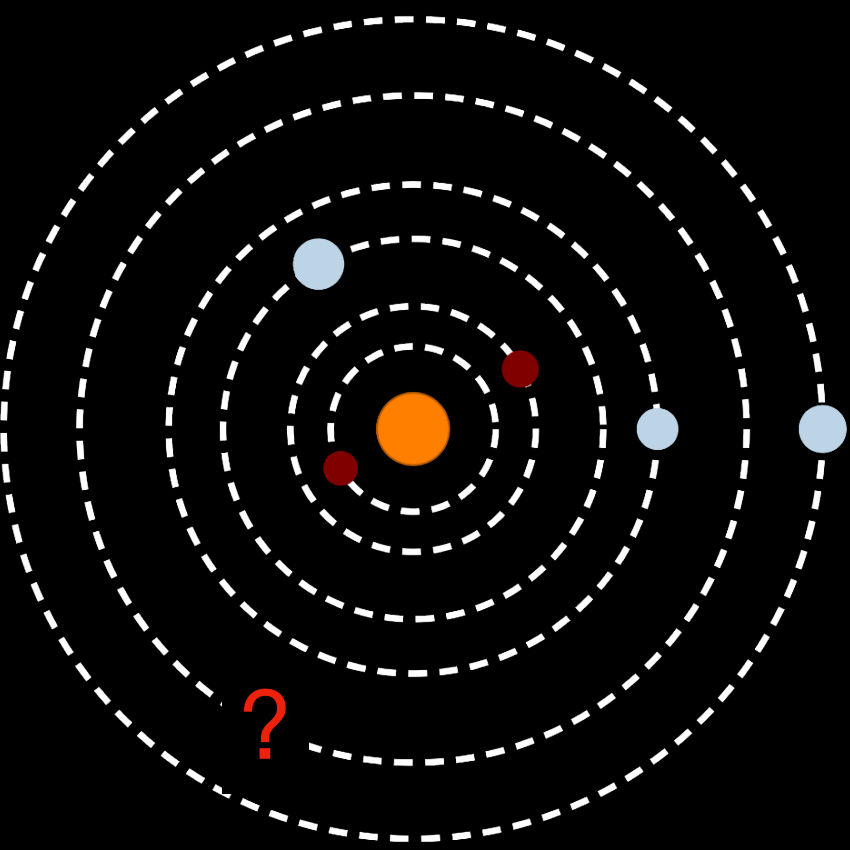
$$\Delta_1 = \Delta_2$$

$$P_{Super,ef} = P_{Super,fg}$$

$$P_{Super,12} = \frac{1}{\Delta_1} = \frac{1}{k_1/P_1 - (k_1 + 1)/P_2}$$



# prediction of the period of TOI-178f



$$P_{Super,cd} = 263d$$

$$P_{Super,de} = 260.5d$$

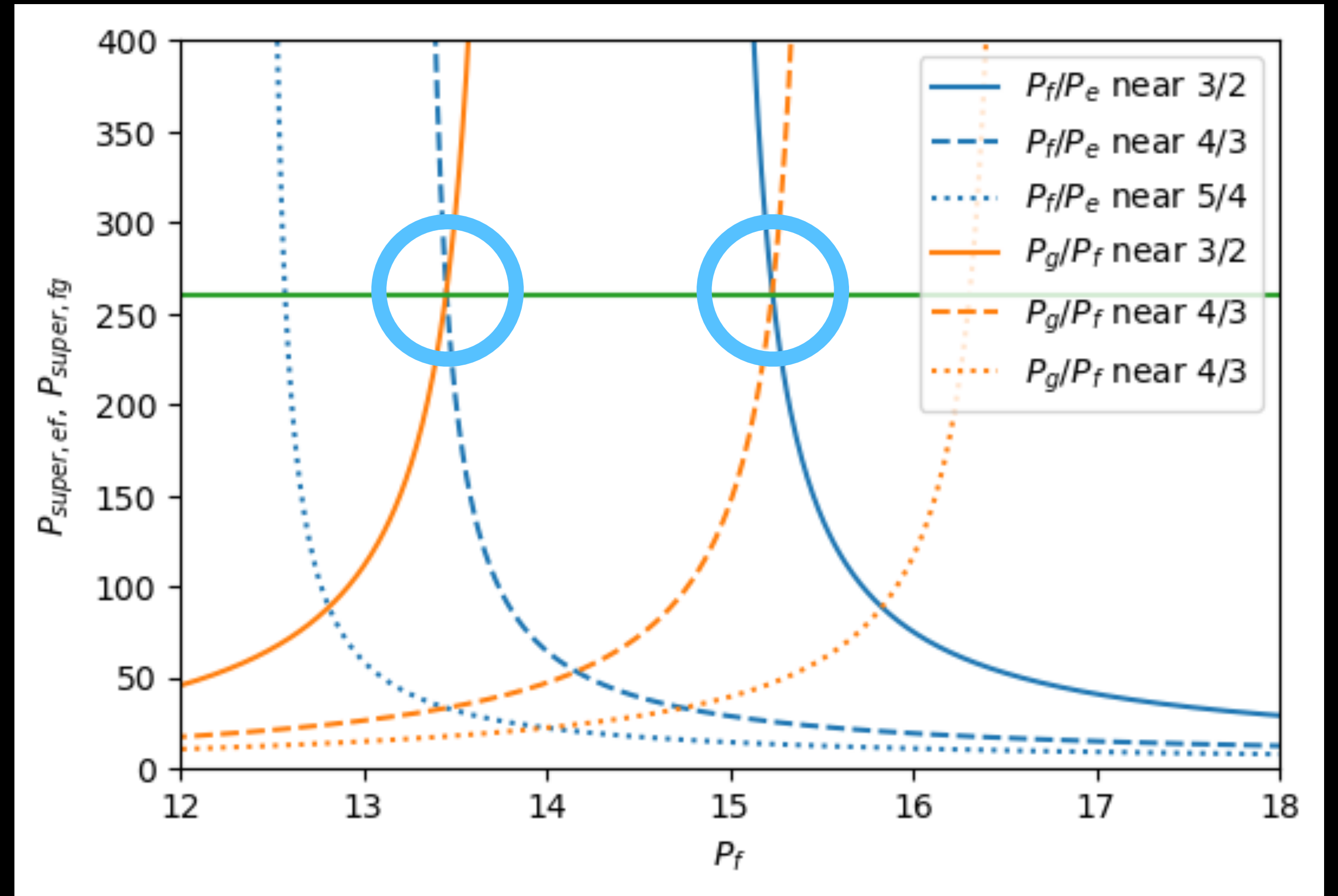
$$P_{Super,eg} = 262.6d$$

Laplace relation if:

$$\Delta_1 = \Delta_2$$

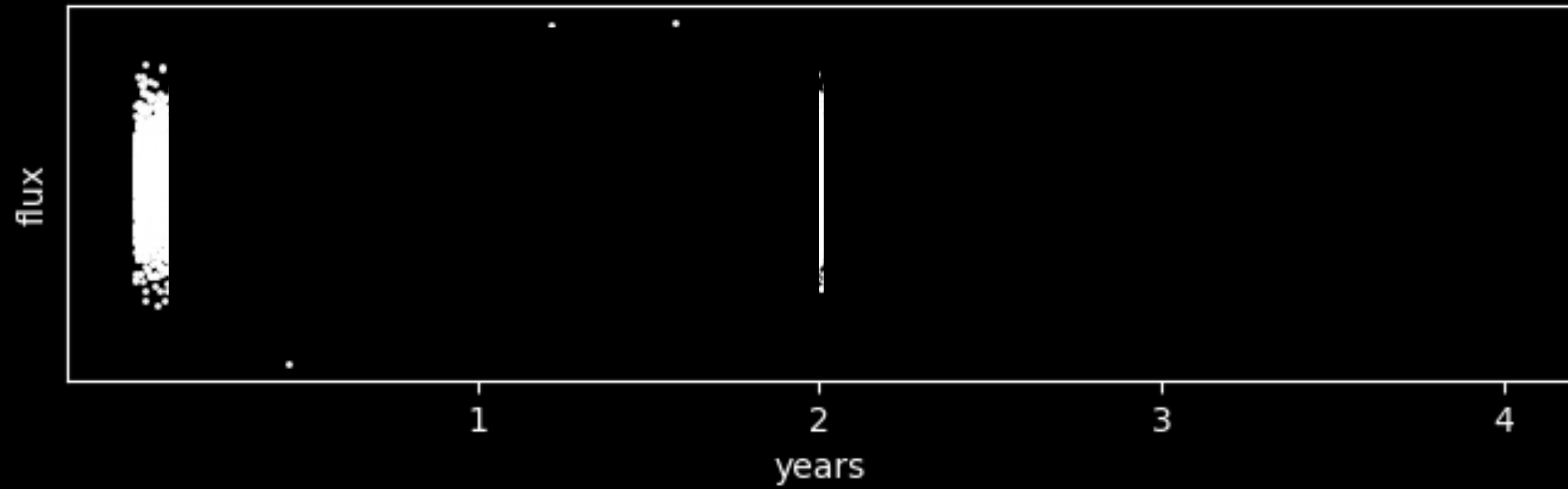
$$P_{Super,ef} = P_{Super,fg}$$

$$P_{Super,12} = \frac{1}{\Delta_1} = \frac{1}{k_1/P_1 - (k_1 + 1)/P_2}$$



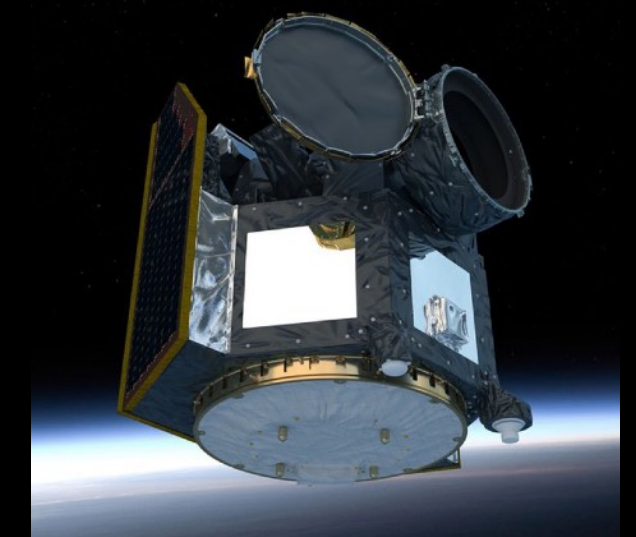
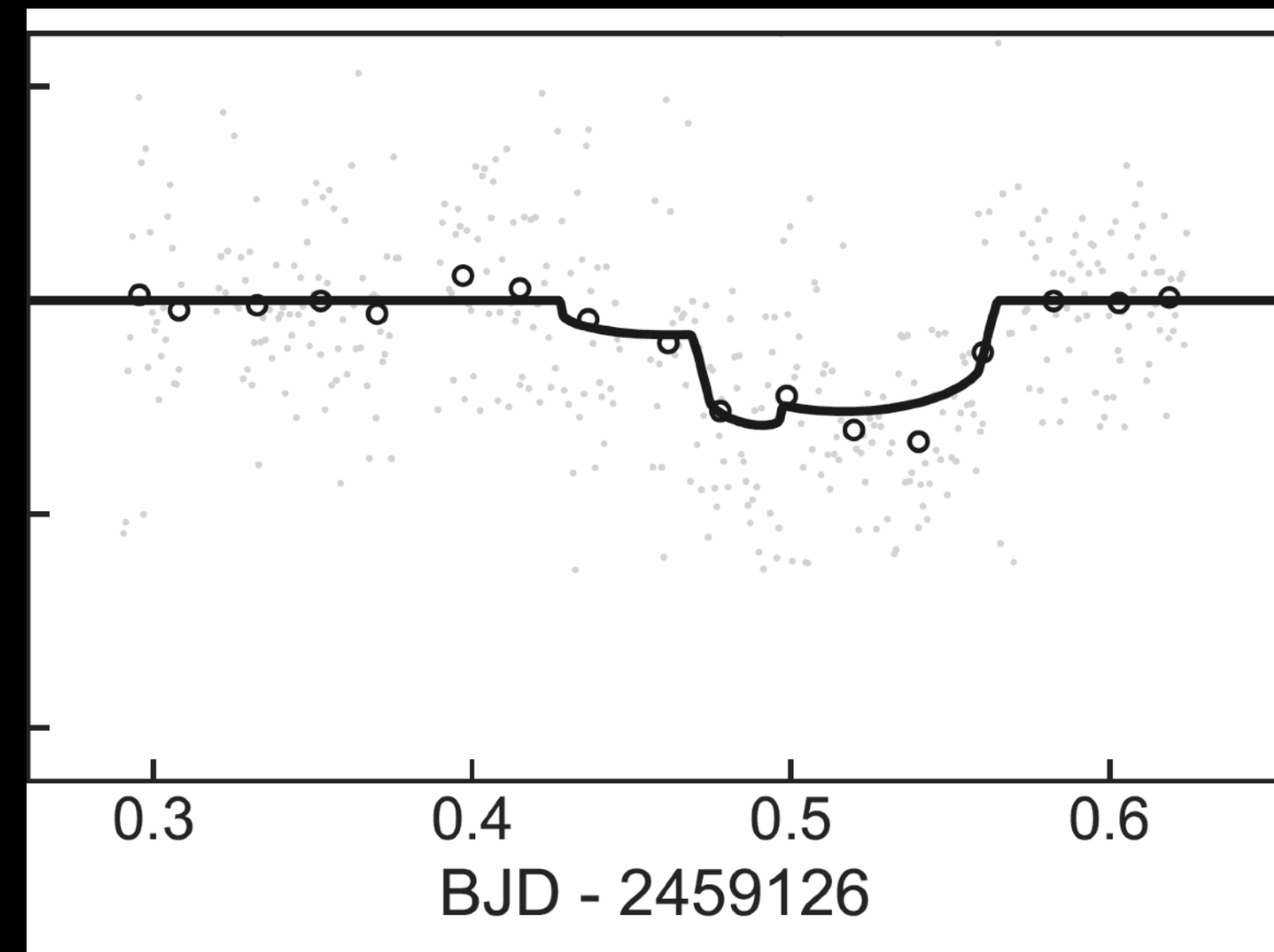
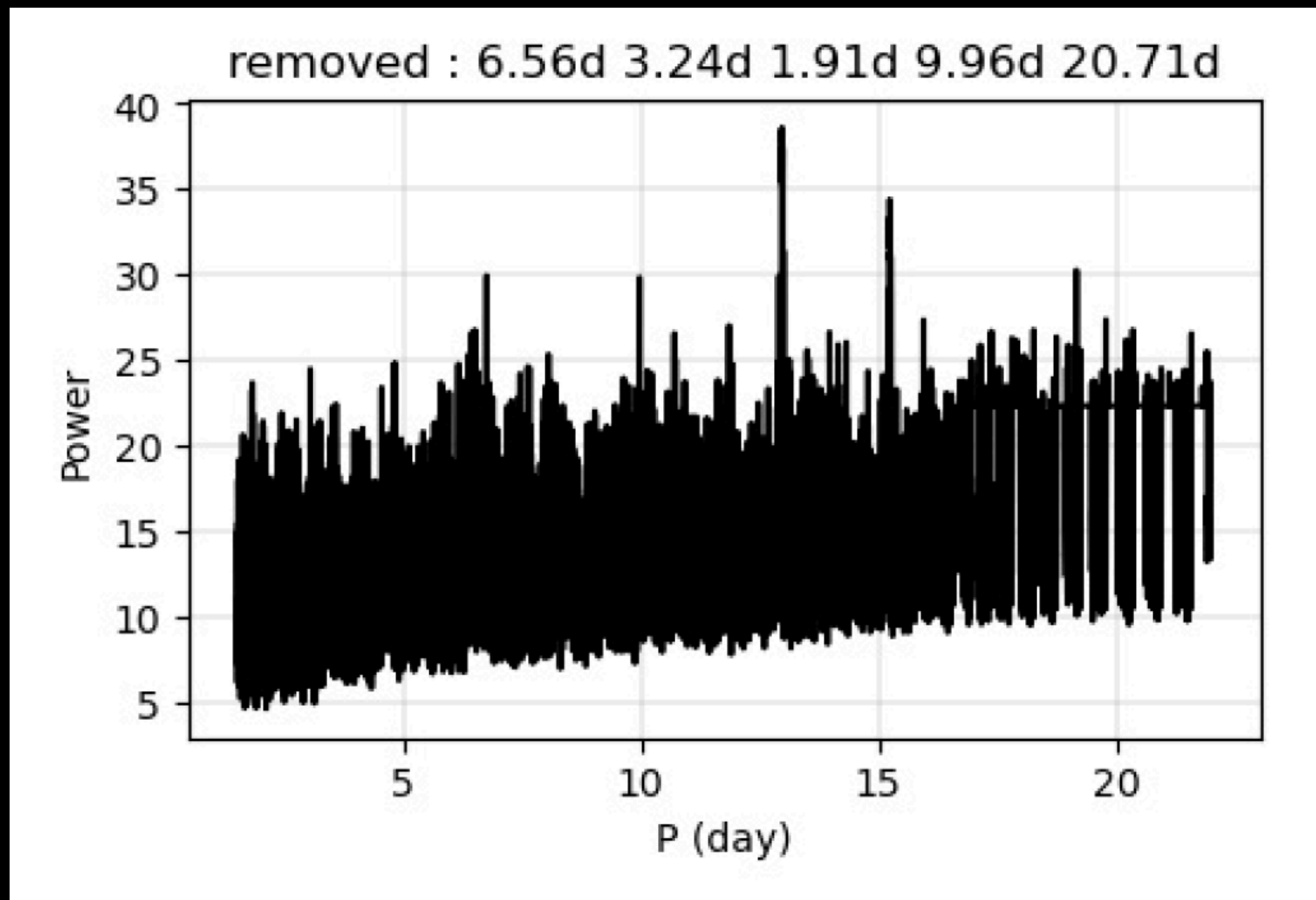
Laplace resonances yield two possible periods:  
 $P = 13.4527$  d, or  $P = 15.2318$  d

# completion of a resonant chain: TOI-178



Laplace resonances yield two possible periods:  
 $P = 13.4527$  d, or  $P = 15.2318$  d

BLS shows peaks at  $\sim 12.9$ d and  $\sim 15.2$ d  
With predicted transits uncertainties of  $\sim 1$  day  
due to a 2 year error propagation

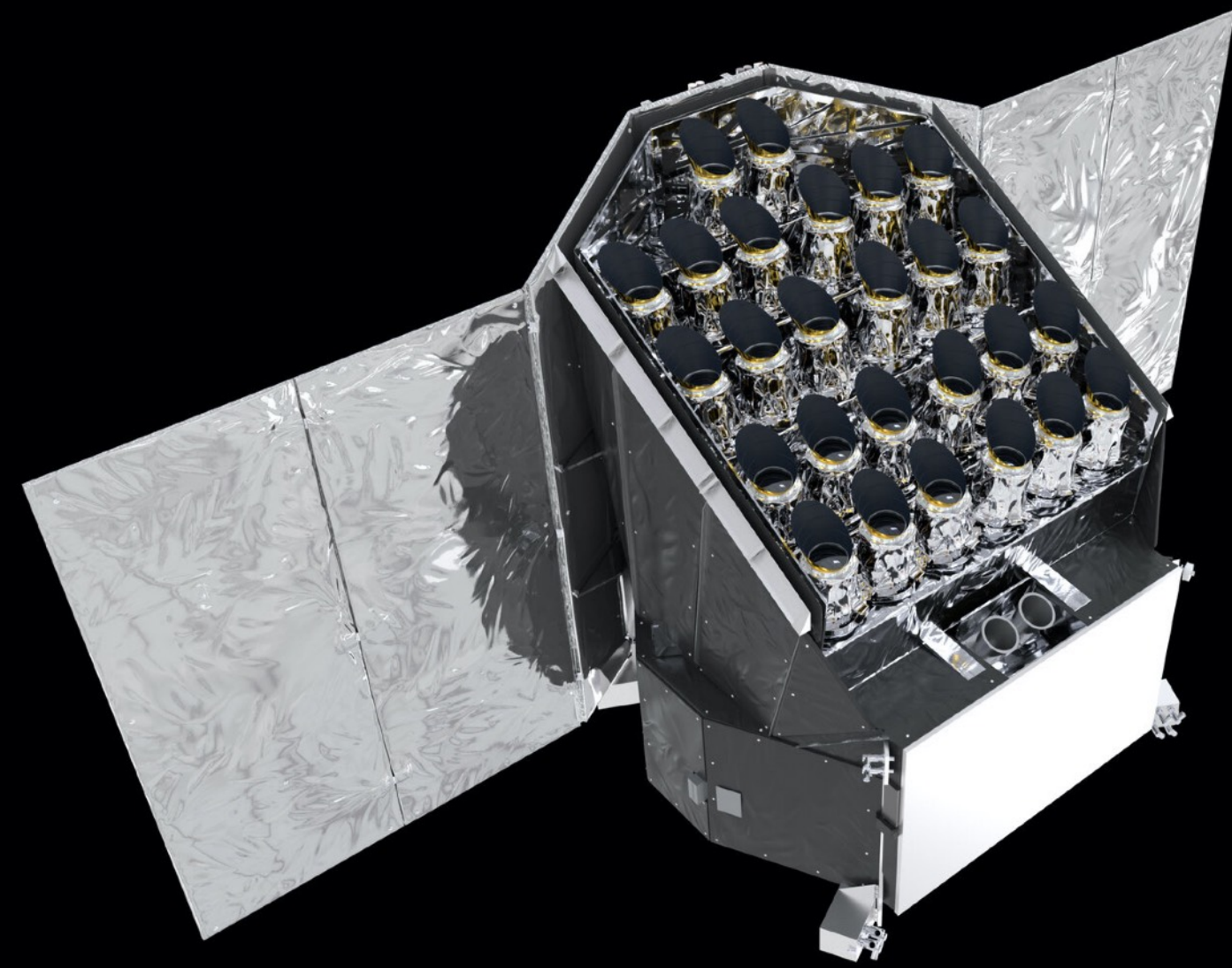


CHEOPS

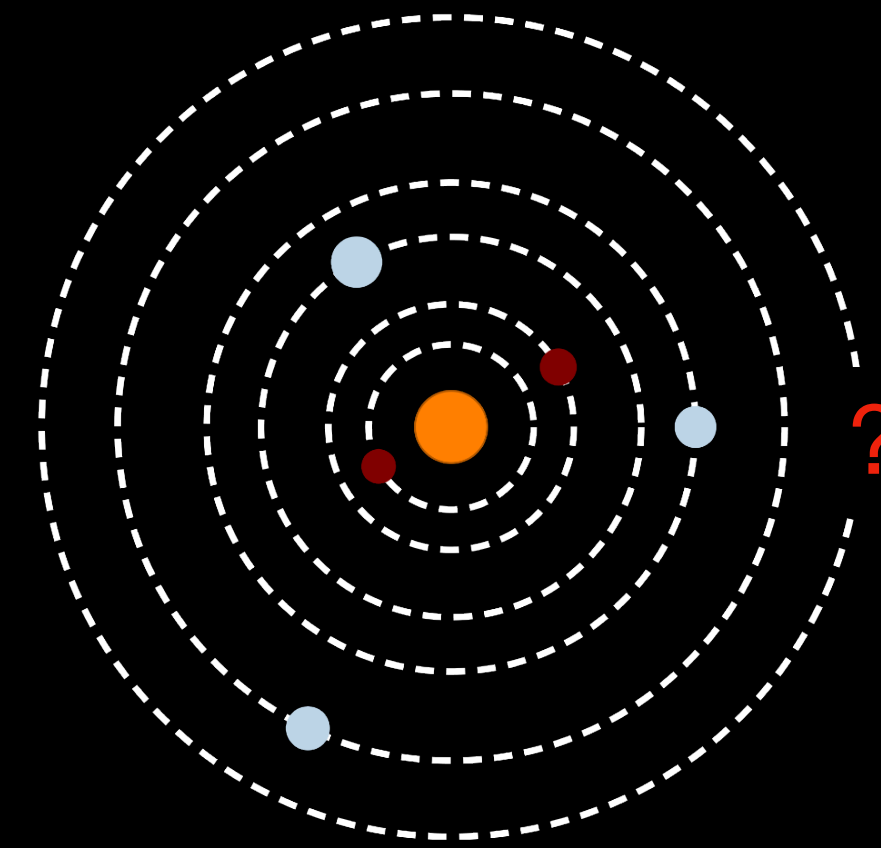
TOI-178 (Leleu+ 2021)

Predicted period of 15.2318 d,  
confirmed period of 15.231915d  $\pm$  0.0001

# Case 2.2



Completing a  
resonant system



Planet continuing the chain,  
single transit



# Generalised Laplace Angle

mean longitude

$$\Psi = l\lambda_1 - (l + m)\lambda_2 + m\lambda_3$$

slowly evolving  $\dot{\lambda} = 2\pi/P$

$$\dot{\Psi} = l/P_1 - (l + m)/P_2 + m/P_3 \approx 0$$

Laplace relation

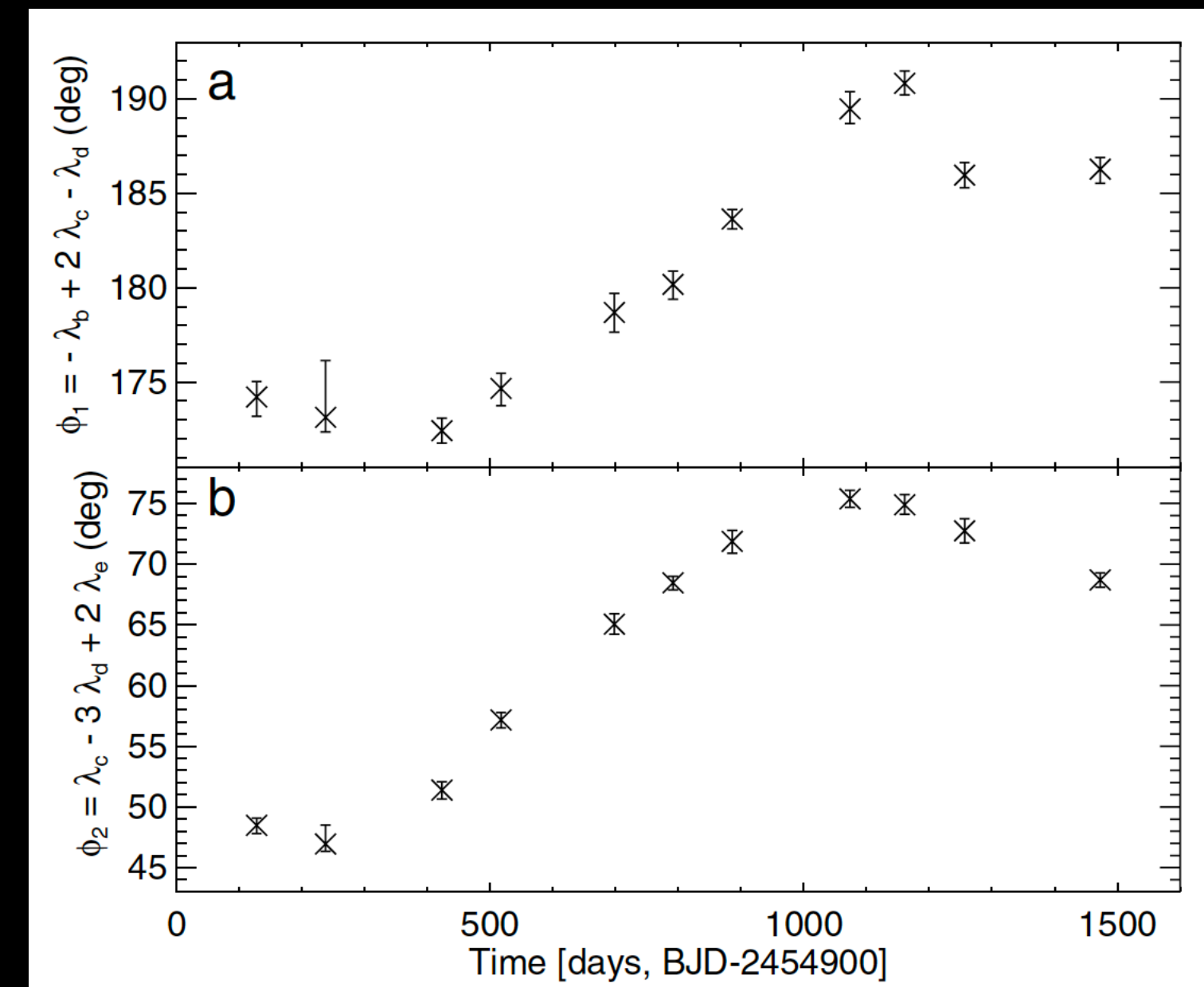
$$\lambda = l - 2e \sin(l - \omega)$$

true longitude

transit occurs when  $l = \pi/2$

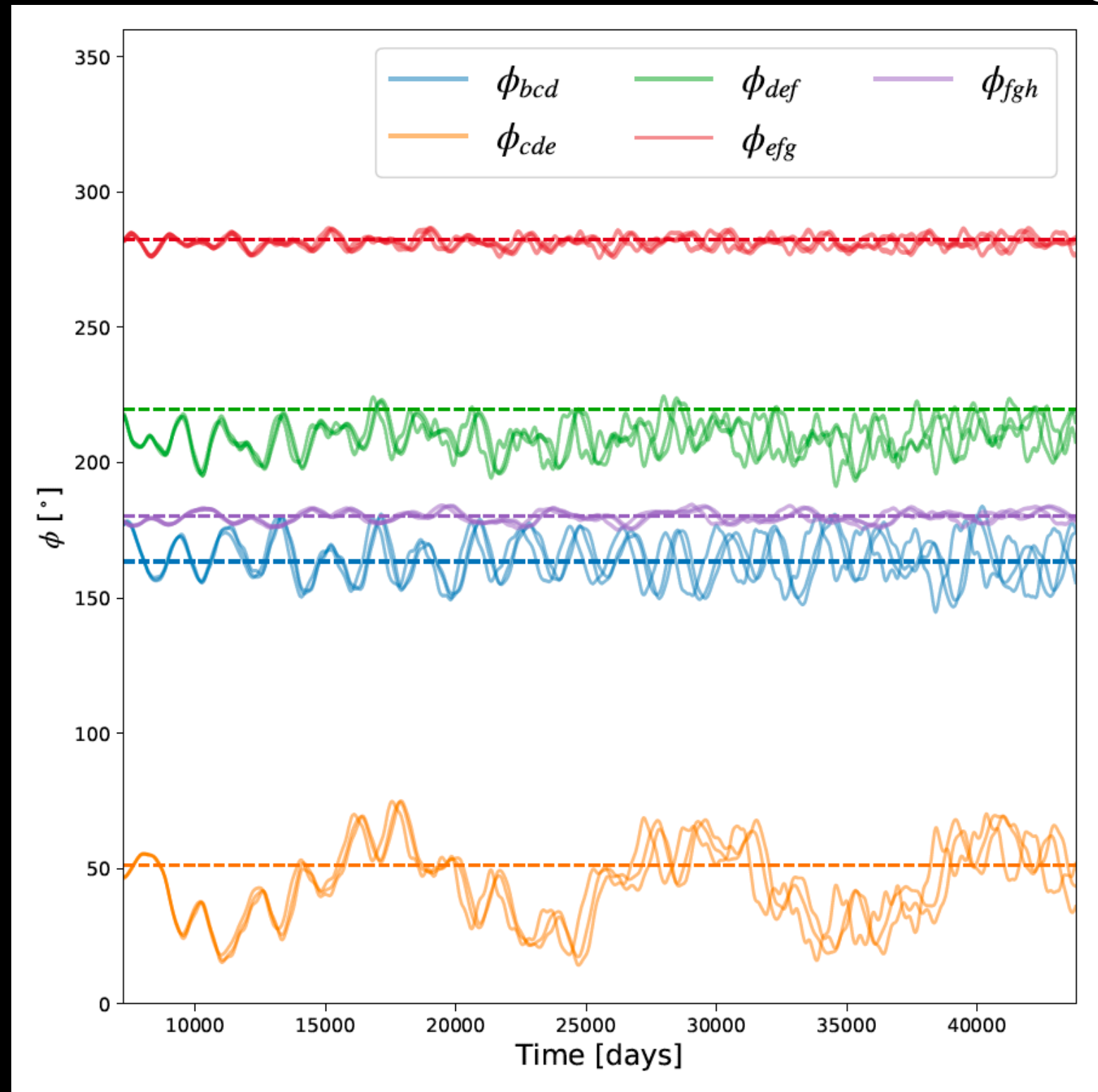
Assuming 0 eccentricities,  $\Psi$  can be computed from [times of transit](#) and [periods](#).

Kepler-223 (Mills+ 2016)

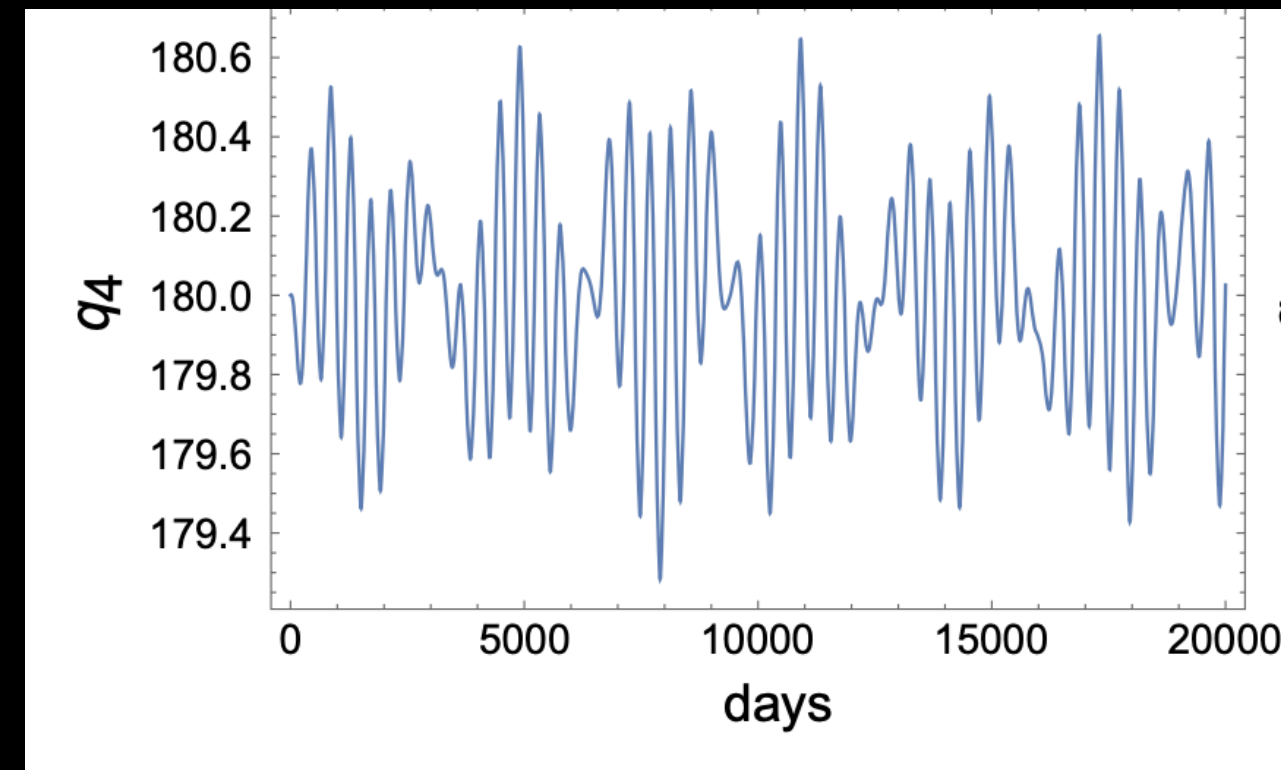


Assuming 0 eccentricities

# Evolution of generalised Laplace angles

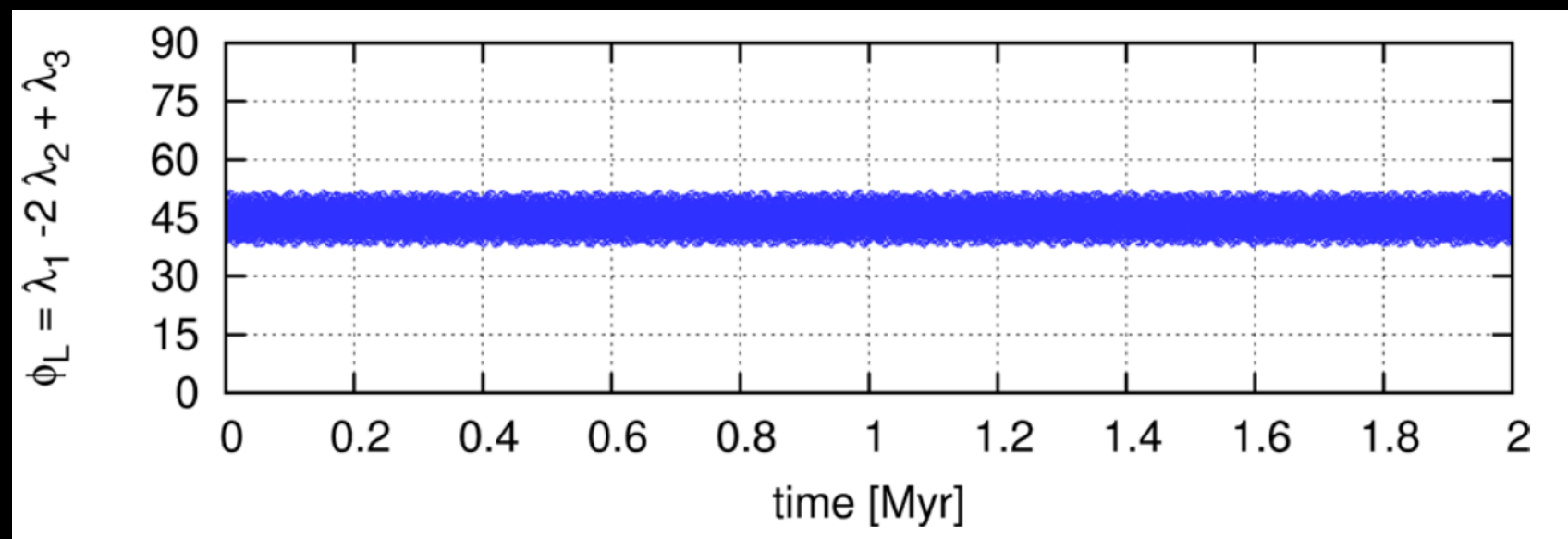


Trappist-1 (Agol 2021)

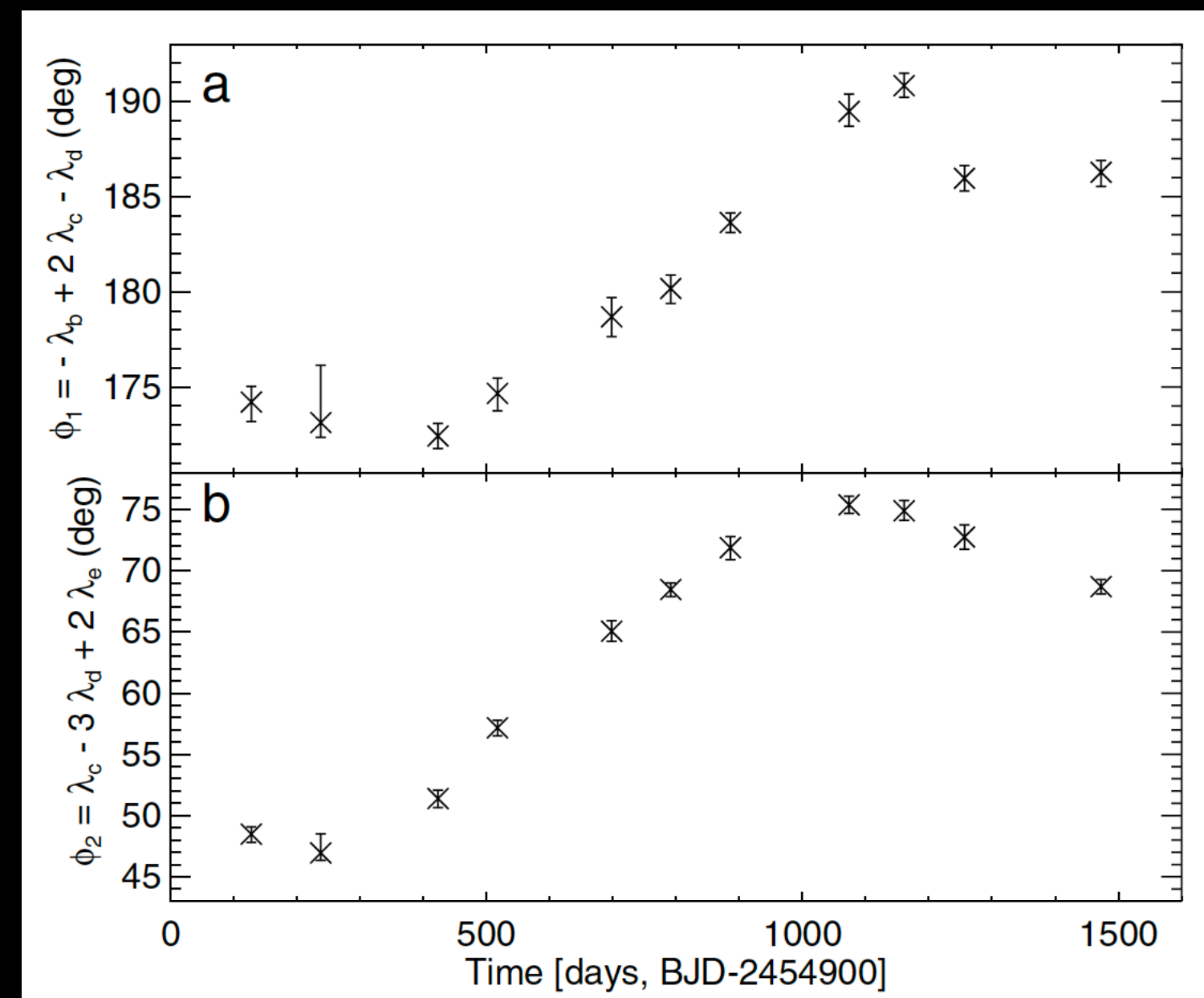


Galilean satellites (Celletti+ 2018)

Typical oscillation  
semi-amplitude  
 $\lesssim 20^\circ$



Kepler-60 (Goździewski 2016)



Kepler- 223 (Mills+ 2016)

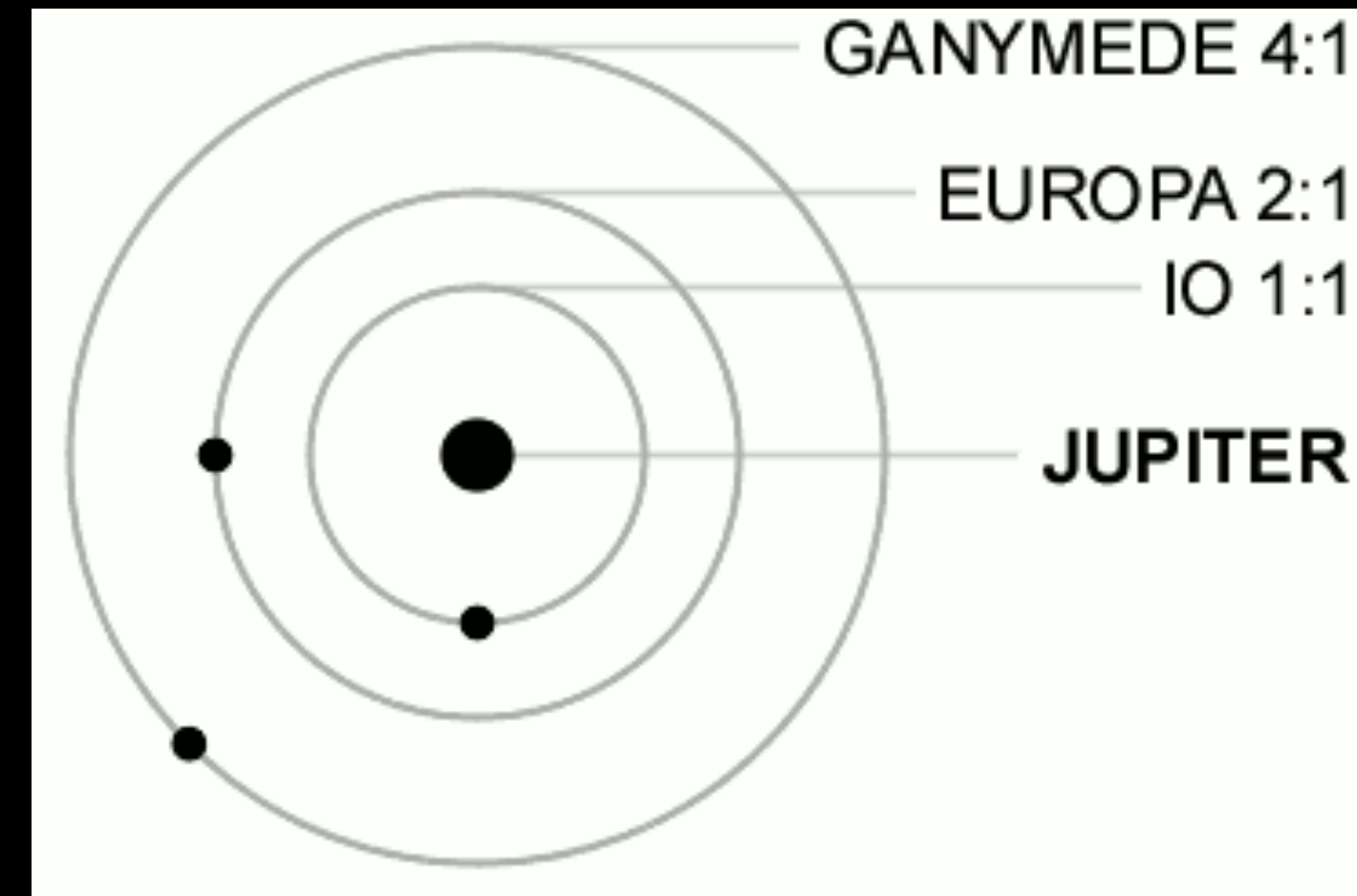
# Laplace (Generalised) Resonant Angle

$$\Psi = l\lambda_1 - (l + m)\lambda_2 + m\lambda_3$$

In the case of a chain of two first-order MMRs

$$\Psi_{eq} = 180^\circ$$

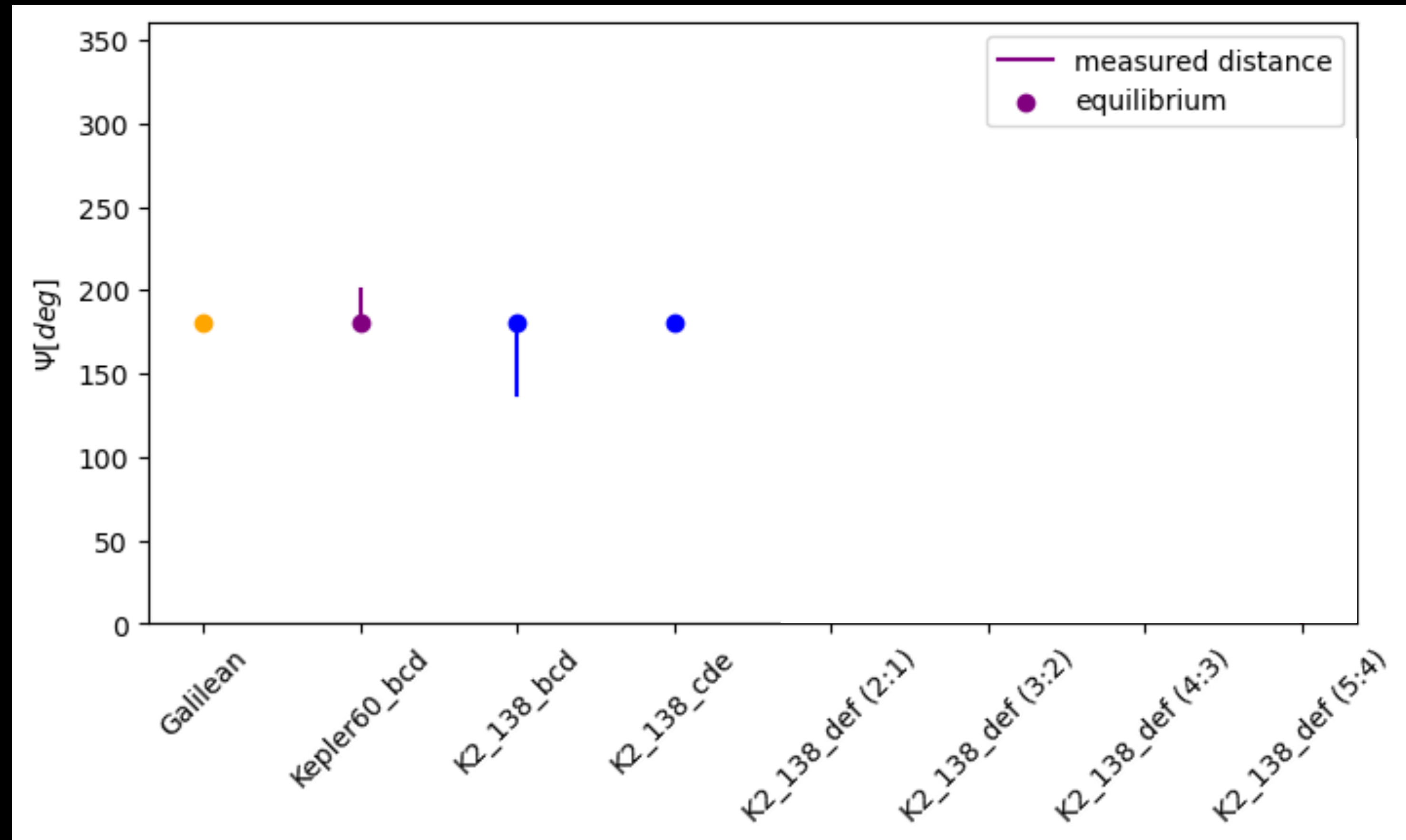
**Sinclair (1975)**



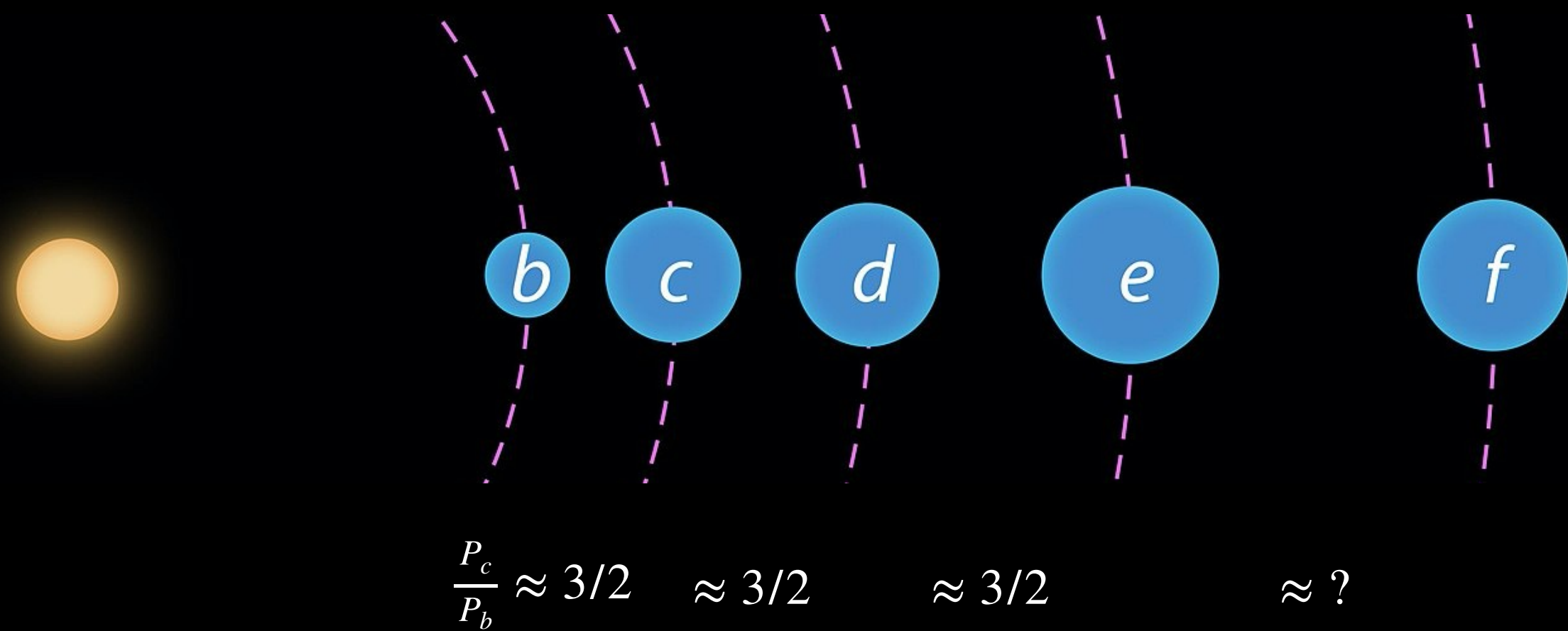
# Distance to equilibrium

Typical oscillation  
semi-amplitude  
 $\lesssim 20^\circ$

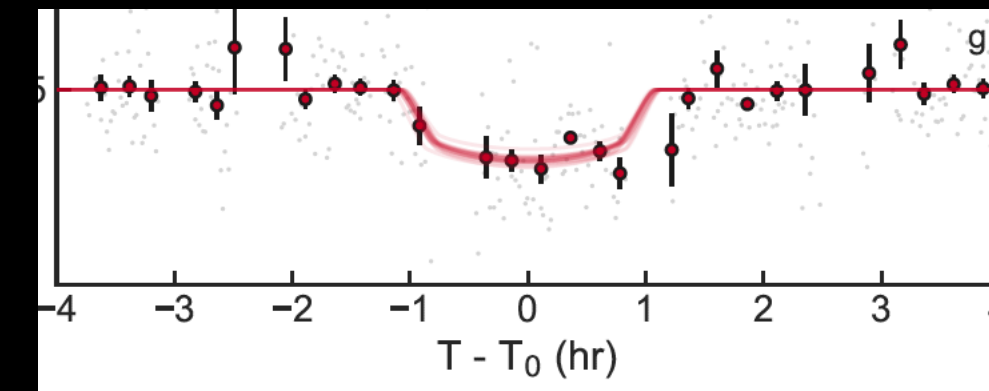
$$\Psi_{eq} = 180^\circ$$



# Where is K2-138f ?



Transit of f - single event



$$P_{Super,ef} = \frac{1}{k_1/P_e - (k_1 + 1)/P_f} = 145$$

$$P_{Super,bc} = 138d$$

$$P_{Super,cd} = 148d$$

$$P_{Super,de} = 145d$$

$$P_f/P_e \approx 2/1 \implies P_f = 15.88d$$

$$P_f/P_e \approx 3/2 \implies P_f = 12.77d$$

$$P_f/P_e \approx 4/3 \implies P_f = 11.23d$$

$$P_f/P_e \approx 5/4 \implies P_f = 10.48d$$

# Where is K2-138?

$$P_{Super,ef} = \frac{1}{k_1/P_e - (k_1 + 1)/P_f}$$

$$P_f/P_e \approx 2/1 \implies P_f = 15.88d$$

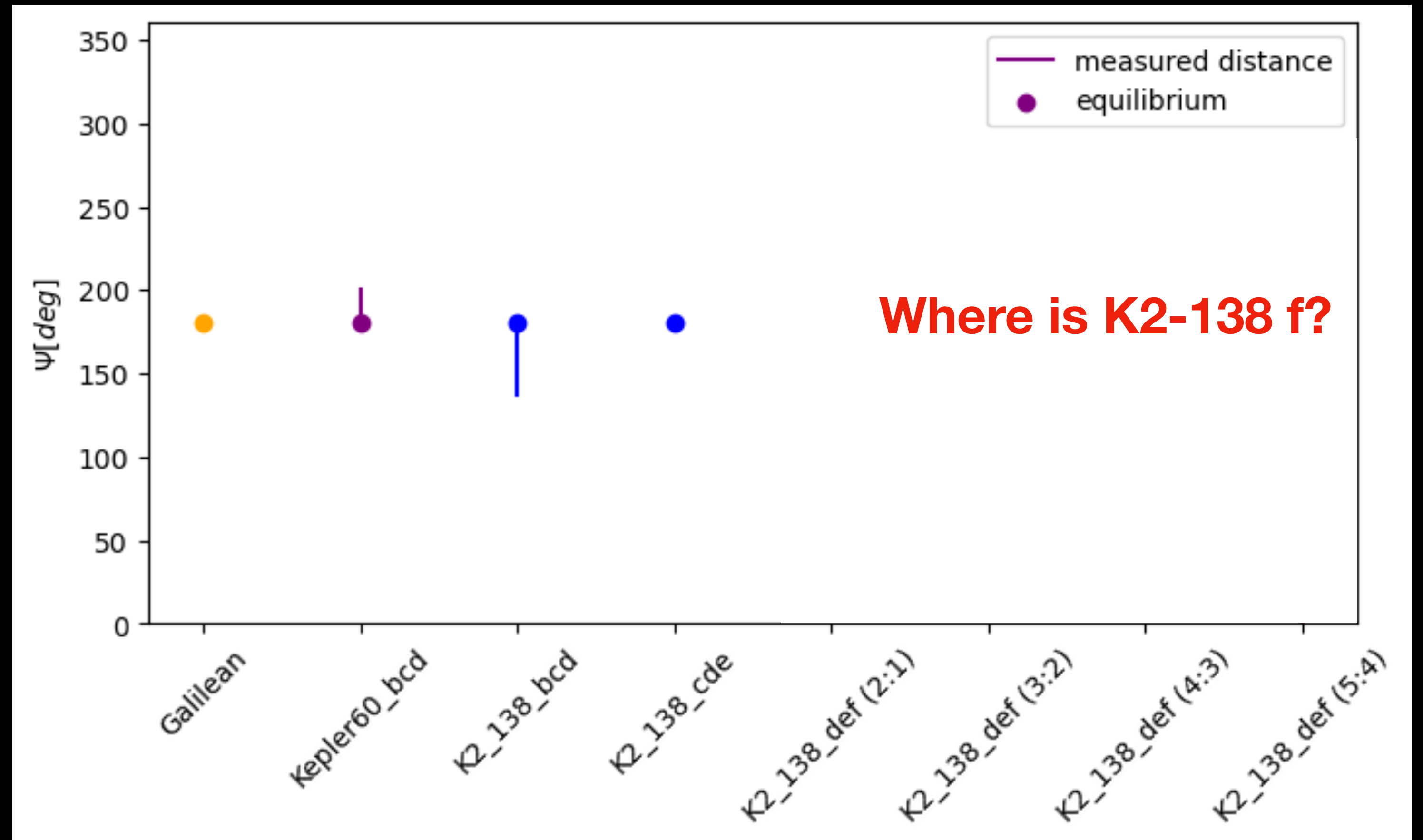
$$P_f/P_e \approx 3/2 \implies P_f = 12.77d$$

$$P_f/P_e \approx 4/3 \implies P_f = 11.23d$$

$$P_f/P_e \approx 5/4 \implies P_f = 10.48d$$

Typical oscillation  
semi-amplitude  
 $\lesssim 20^\circ$

$$\Psi_{eq} = 180^\circ$$



# Distance to equilibrium

$$P_{Super,ef} = \frac{1}{k_1/P_e - (k_1 + 1)/P_f}$$

$$P_f/P_e \approx 2/1 \implies P_f = 15.88d$$

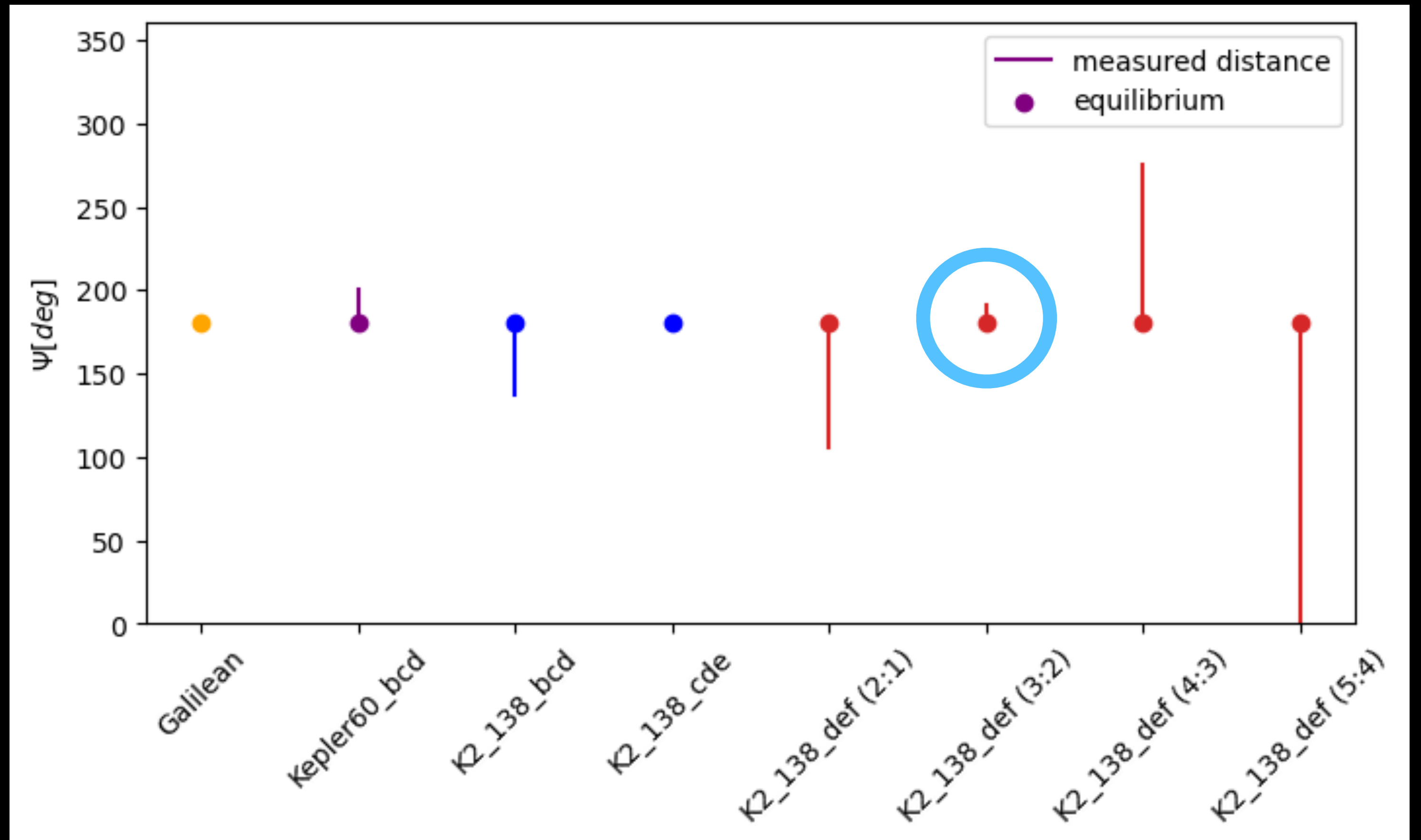
$$P_f/P_e \approx 3/2 \implies P_f = 12.77d$$

$$P_f/P_e \approx 4/3 \implies P_f = 11.23d$$

$$P_f/P_e \approx 5/4 \implies P_f = 10.48d$$

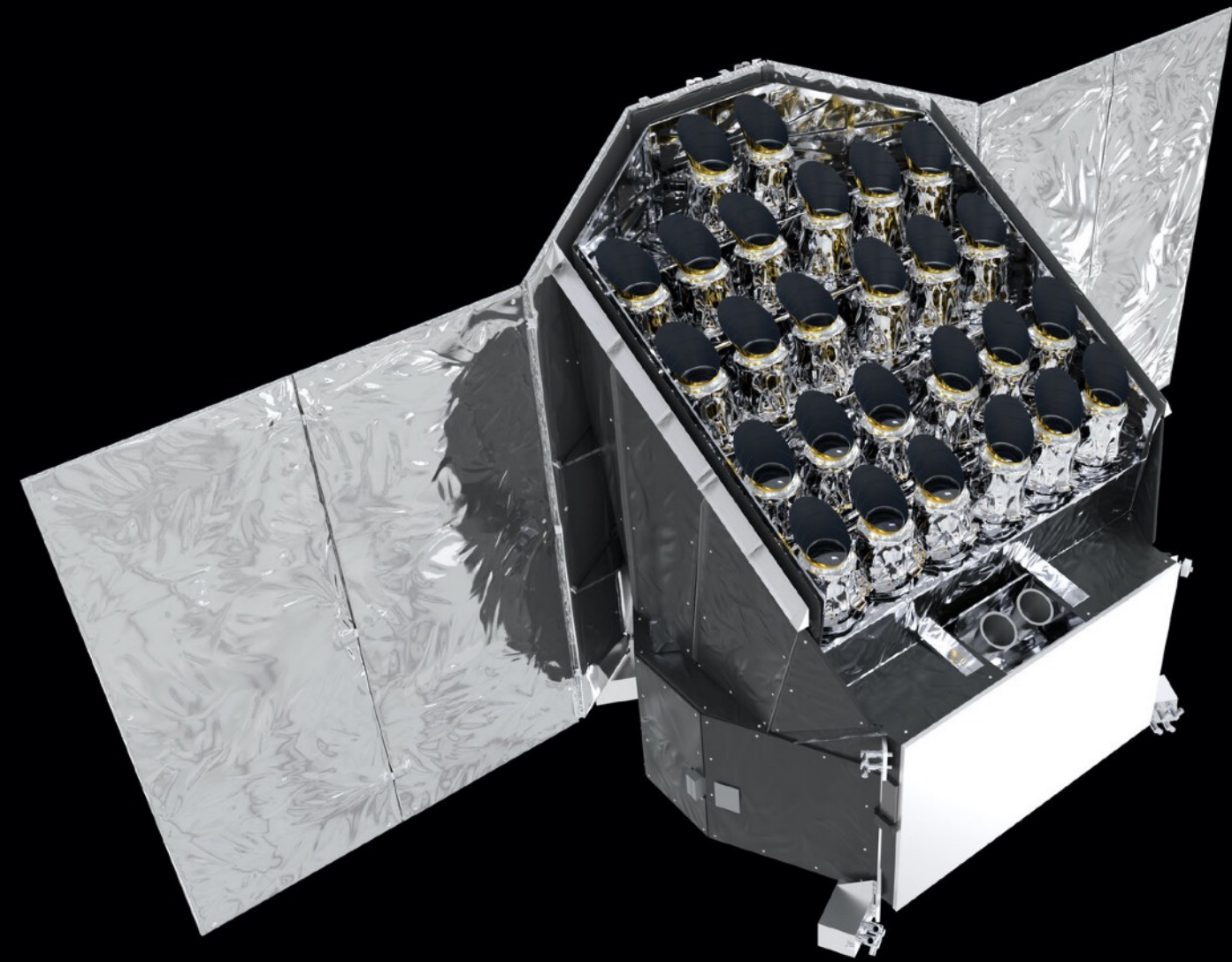
Typical oscillation  
semi-amplitude  
 $\lesssim 20^\circ$

$$\Psi_{eq} = 180^\circ$$

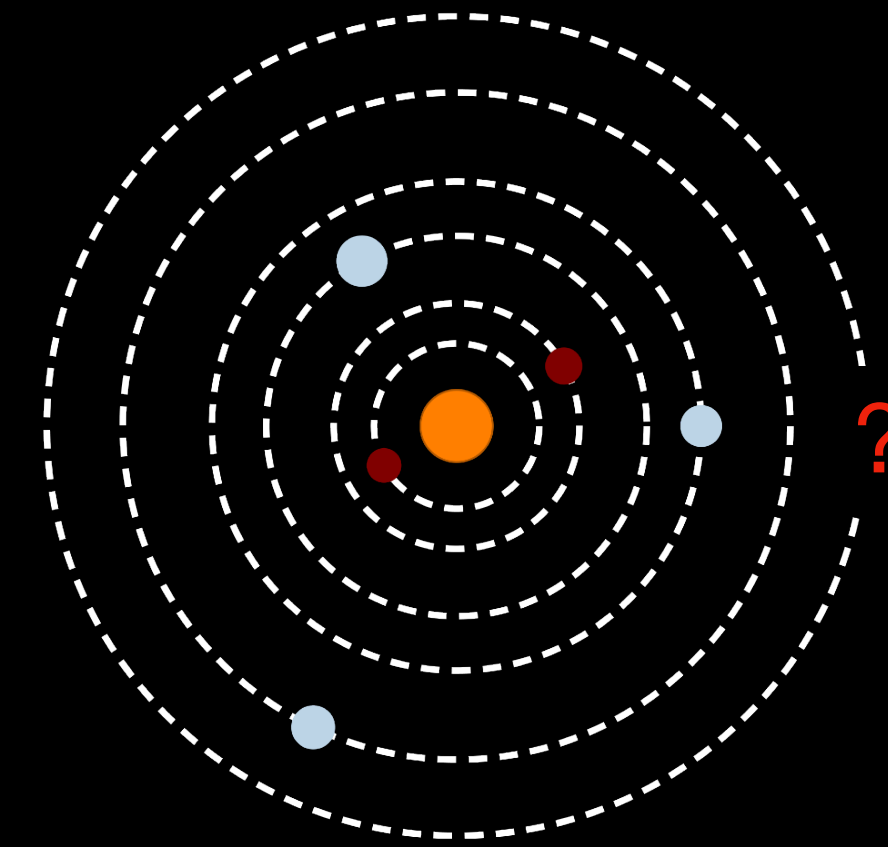


# Case 2.3

Continuous observation /  
Step and stare



Completing a  
resonant system

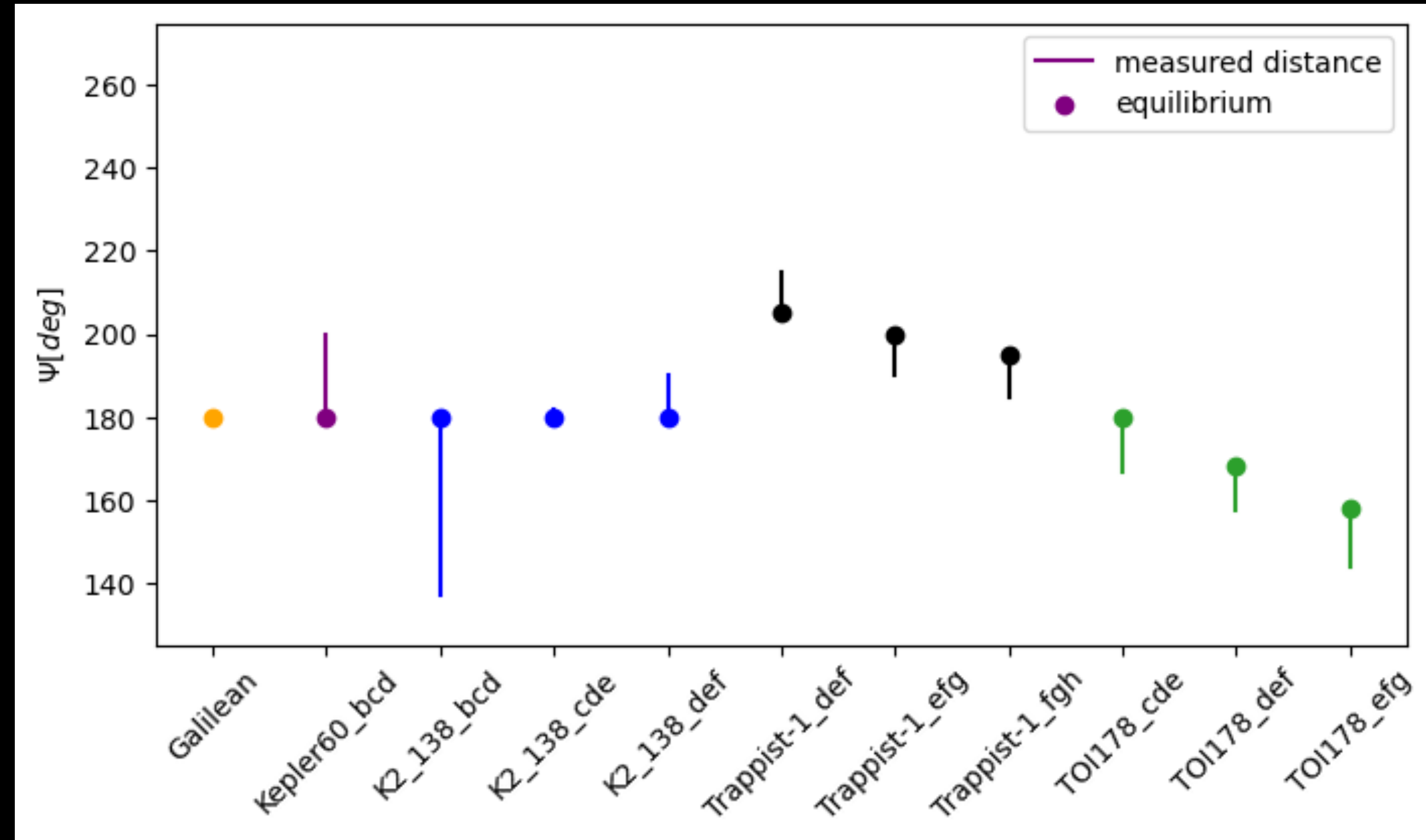


Planet continuing the chain,  
no transit

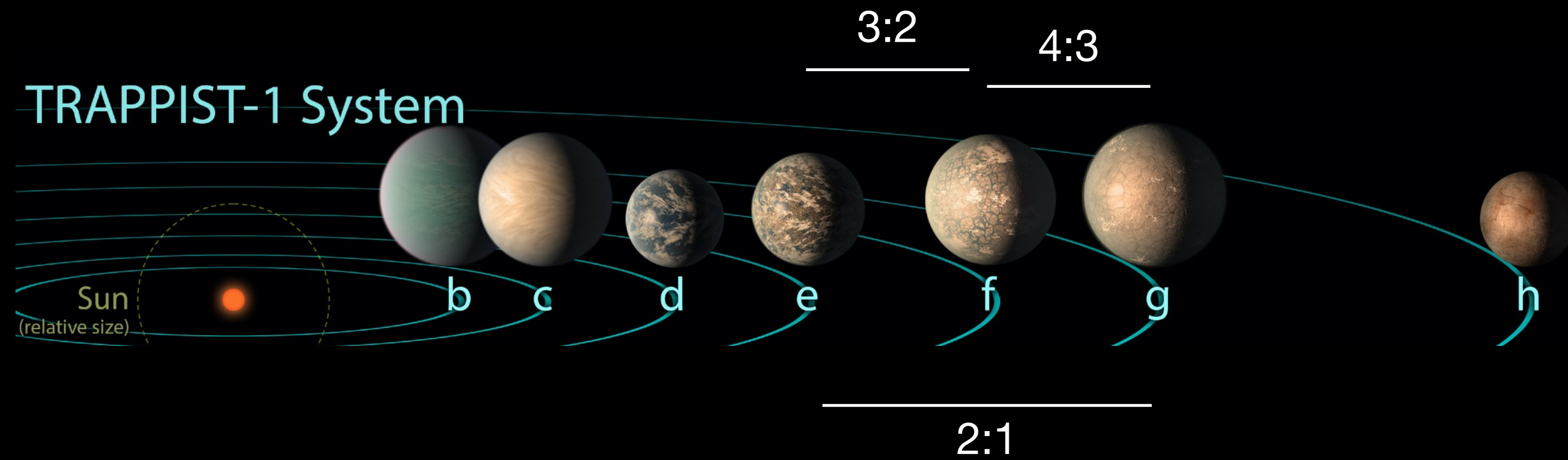


# Distance to equilibrium - general case

Typical oscillation  
semi-amplitude  
 $\lesssim 20^\circ$



# Trappist-1



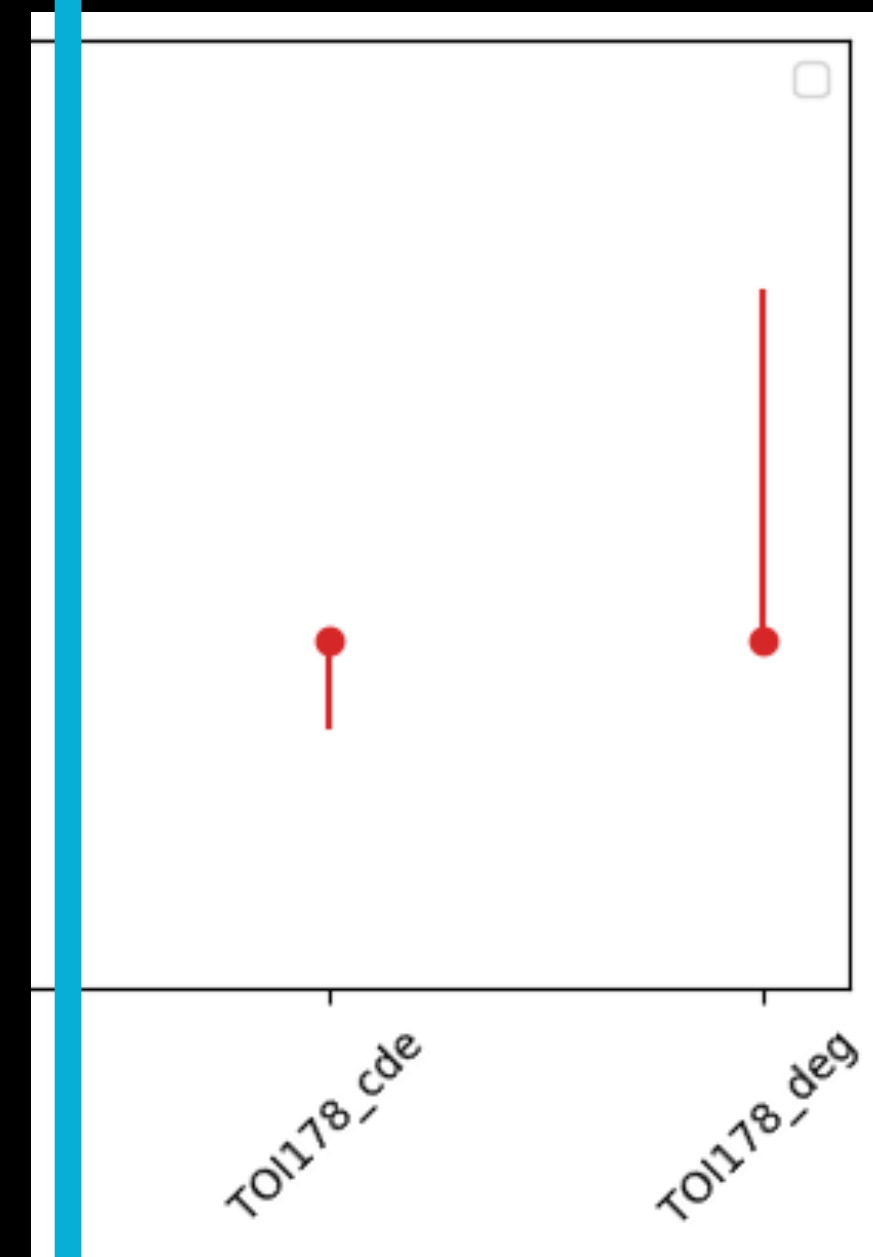
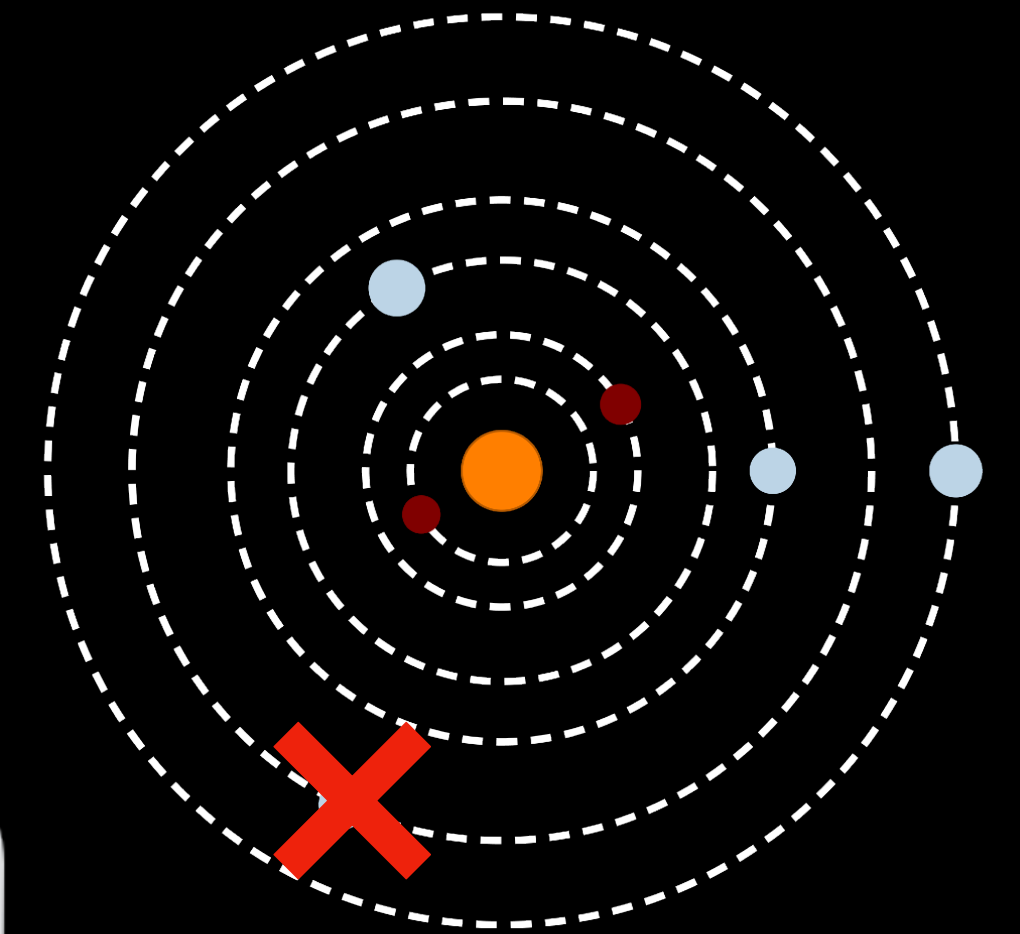
Trappist:

$$P_f/P_e \approx 3/2, P_g/P_f \approx 4/3$$
$$\implies P_g/P_e \approx 2$$

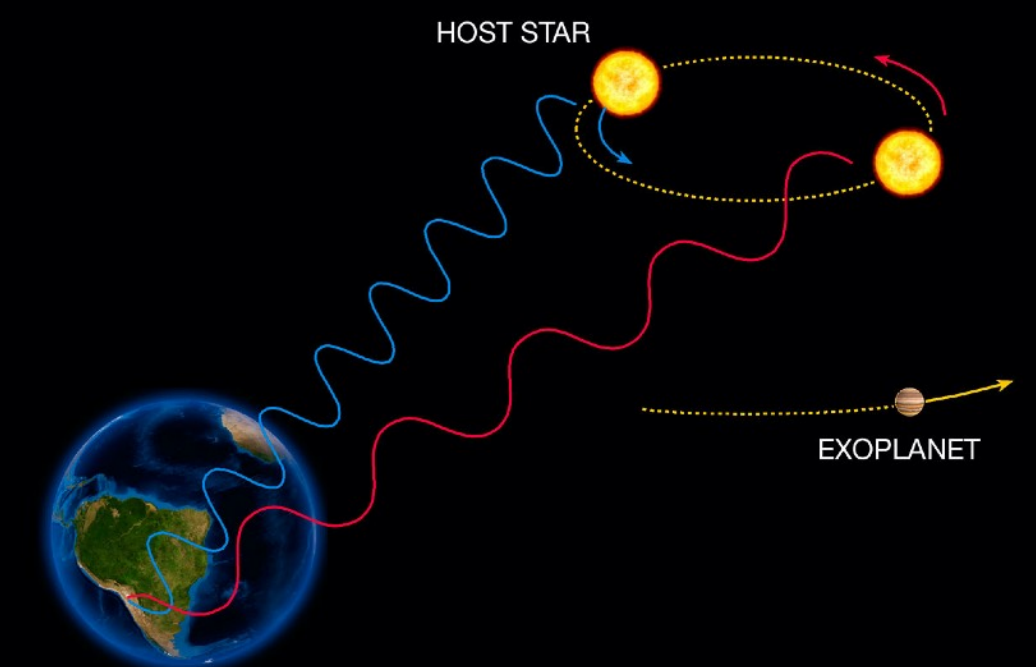
Computation of the fixed points Delisle (2017)

# TOI-178 Missing planet

TOI-178 with 5 planets

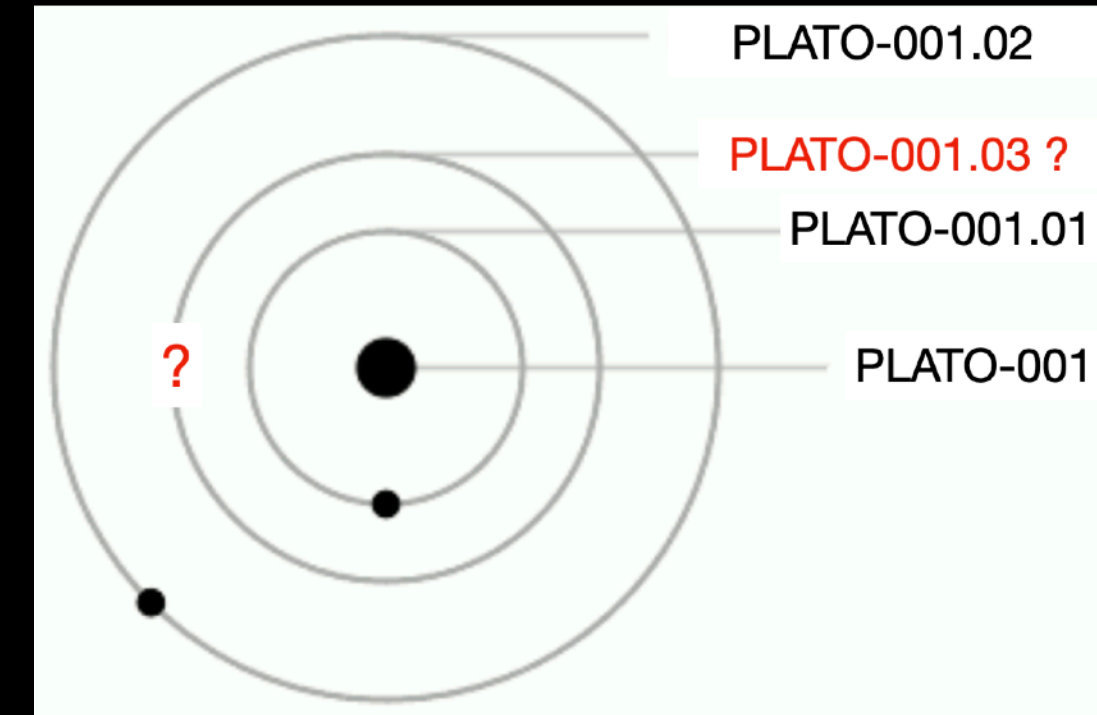


Radial velocity follow-up

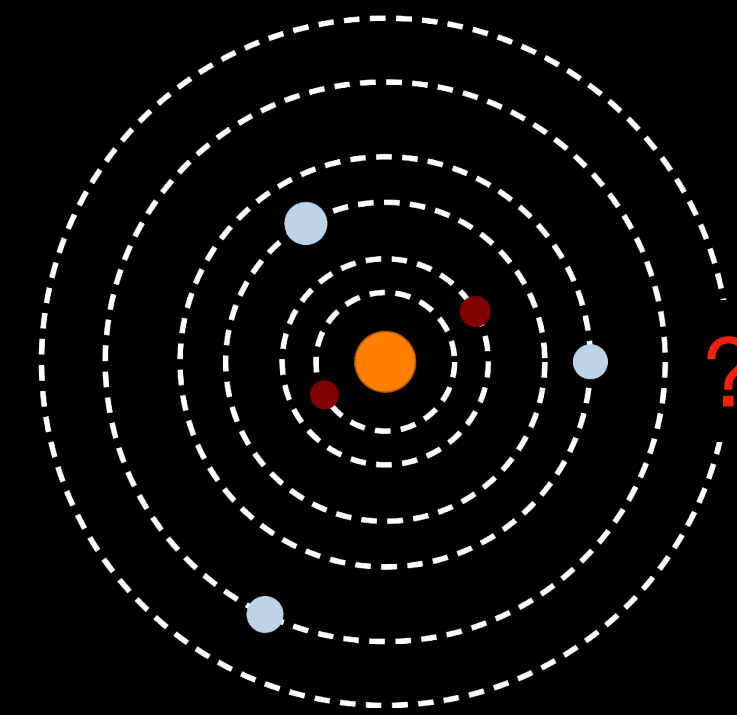


# Conclusions

Atypical mean-motion resonances can hint at planets



In case of known chain, a single transit is enough to predict the period of an additional planet



The hint of a totally missed planet can be recognised by odd Laplace angles

